

Magic Baseline Beta Beam

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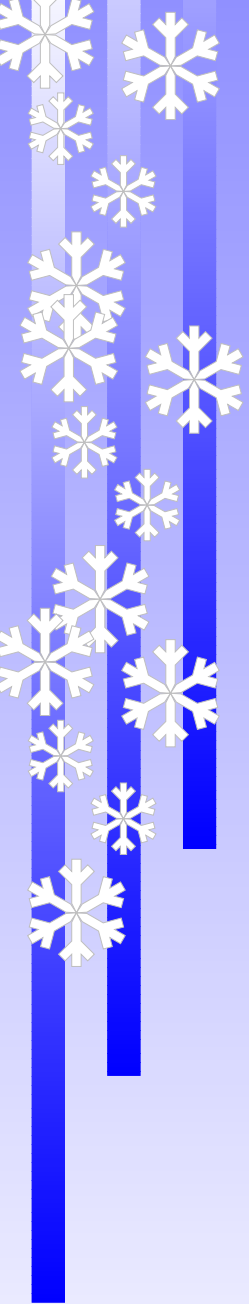
work done in collaboration with

S. Choubey, A. Raychaudhuri and A. Samanta

Phys.Lett.B629(2005)33-40

Nucl.Phys.B771(2007)1-27

PLAN



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- Present Status & Missing Links
- “Golden Channel”, $P_{e\mu}$

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- CERN based β -beam source
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- India-Based Neutrino Observatory (INO)
- Results
- Conclusions

Then and Now

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Now (2004): APS Neutrino Study Group

Much of what we know about neutrinos we have learned in the last six years. . . . We have so many new questions, . . . We are most certain of one thing : neutrinos will continue to surprise us



Past and Present

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Next generation experiments have been planned/proposed world-wide to further pin down the values of the oscillation parameters

Present Status

⇒ **ATM + K2K + MINOS :**

$$|\Delta m_{31}^2| \simeq 2.5_{-0.5}^{+0.7} \times 10^{-3} \text{eV}^2, \theta_{23} \simeq 45.0^{\circ}_{-9.22^{\circ}}^{+9.22^{\circ}} \quad (3\sigma)$$

⇒ **Solar Neutrino Experiments + KamLAND :**

$$\Delta m_{21}^2 \simeq 8.0_{-0.8}^{+1.2} \times 10^{-5} \text{eV}^2, \theta_{12} \simeq 33.83^{\circ}_{-3.83^{\circ}}^{+4.82^{\circ}} \quad (3\sigma)$$

(Our convention : $\Delta m_{ij}^2 = m_i^2 - m_j^2$)

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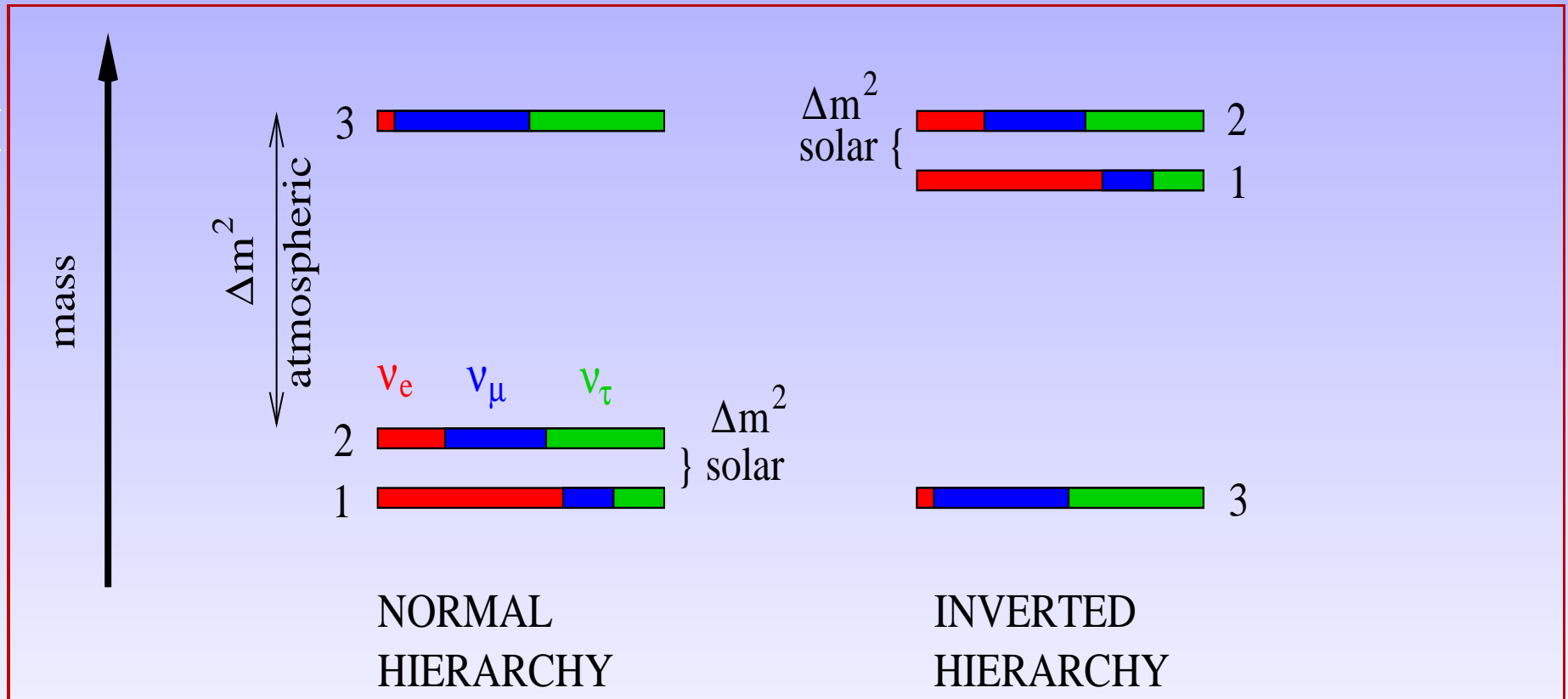
⇒ **Chooz + Palo Verde :**

$$\sin^2 2\theta_{13} < \mathbf{0.17} \quad (3\sigma)$$

⇒ **Two large mixing** angles and the relative oscillation frequencies open the possibility to test CP violation in the neutrino sector, if θ_{13} and δ_{CP} are not vanishingly small

Unsolved Issues

- The sign of Δm_{31}^2 ($m_3^2 - m_1^2$) is not known. Neutrino mass spectrum can be normal or inverted hierarchical



Continued..

- Only an upper limit on $\sin^2 2\theta_{13}$ (< 0.17 at 3σ) exists
- The Dirac CP phase (δ_{CP}) is unconstrained

We will focus on the first two issues...

Golden Channel ($P_{e\mu}$)

The appearance probability ($\nu_e \rightarrow \nu_\mu$) in matter, upto second order in the small parameters $\alpha \equiv \Delta m_{21}^2 / \Delta m_{31}^2$ and $\sin 2\theta_{13}$,

$$\begin{aligned} P_{e\mu} \simeq & \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2[(1 - \hat{A})\Delta]}{(1 - \hat{A})^2} \\ & + \alpha \sin 2\theta_{13} \xi \sin \delta_{CP} \sin(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \\ & + \alpha \sin 2\theta_{13} \xi \cos \delta_{CP} \cos(\Delta) \frac{\sin(\hat{A}\Delta)}{\hat{A}} \frac{\sin[(1 - \hat{A})\Delta]}{(1 - \hat{A})} \\ & + \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}; \end{aligned}$$

where $\Delta \equiv \Delta m_{31}^2 L / (4E)$, $\xi \equiv \cos \theta_{13} \sin 2\theta_{21} \sin 2\theta_{23}$,
and $\hat{A} \equiv \pm(2\sqrt{2}G_F n_e E) / \Delta m_{31}^2$

Cervera et al., hep-ph/0002108

Freund, Huber, Lindner, hep-ph/0105071

Eight-fold Degeneracy

■ $(\theta_{13}, \delta_{CP})$ intrinsic degeneracy

Burguet-Castell, Gavela, Gomez-Cadenas, Hernandez, Mena,
hep-ph/0103258

■ $(\text{sgn}(\Delta m_{31}^2), \delta_{CP})$ degeneracy

Minakata, Nunokawa, hep-ph/0108085

■ $(\theta_{23}, \pi/2 - \theta_{23})$ degeneracy

Fogli, Lisi, hep-ph/9604415

Severely deteriorates the sensitivity



Problem & Solution

Degeneracies create “Clone” Solutions

Barger, Marfatia, Whisnant, hep-ph/0112119



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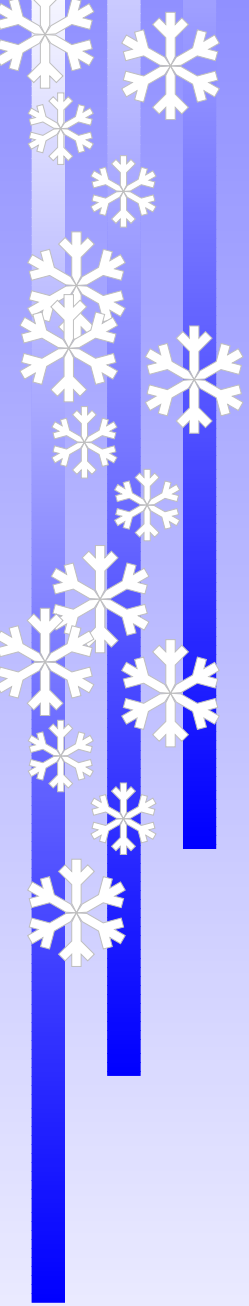
Barger, Marfatia, Whisnant, hep-ph/0112119

Kill the “Clones” at the “Magic” Baseline

Huber, Winter, hep-ph/0301257

Smirnov, hep-ph/0610198

Magic Baseline



Magic Baseline

If one chooses : $\sin(\hat{A}\Delta) = 0$

- The δ_{CP} dependence disappears from $P_{e\mu}$
- Golden channel enables a clean determination of θ_{13} and $\text{sgn}(\Delta m_{31}^2)$

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First non-trivial solution: $\sqrt{2}G_F n_e L = 2\pi$ (indep of E)

- Isoscalar medium of constant density ρ :
 $L_{\text{magic}}[\text{km}] \approx 32725/\rho[\text{gm}/\text{cm}^3]$
- According to PREM, the “Magic Baseline”

$$L_{\text{magic}} = 7690 \text{ km}$$

Best-fit values

$$|\Delta m_{31}^2| = 2.5 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} = 1.0$$

$$\Delta m_{21}^2 = 8.0 \times 10^{-5} \text{ eV}^2$$

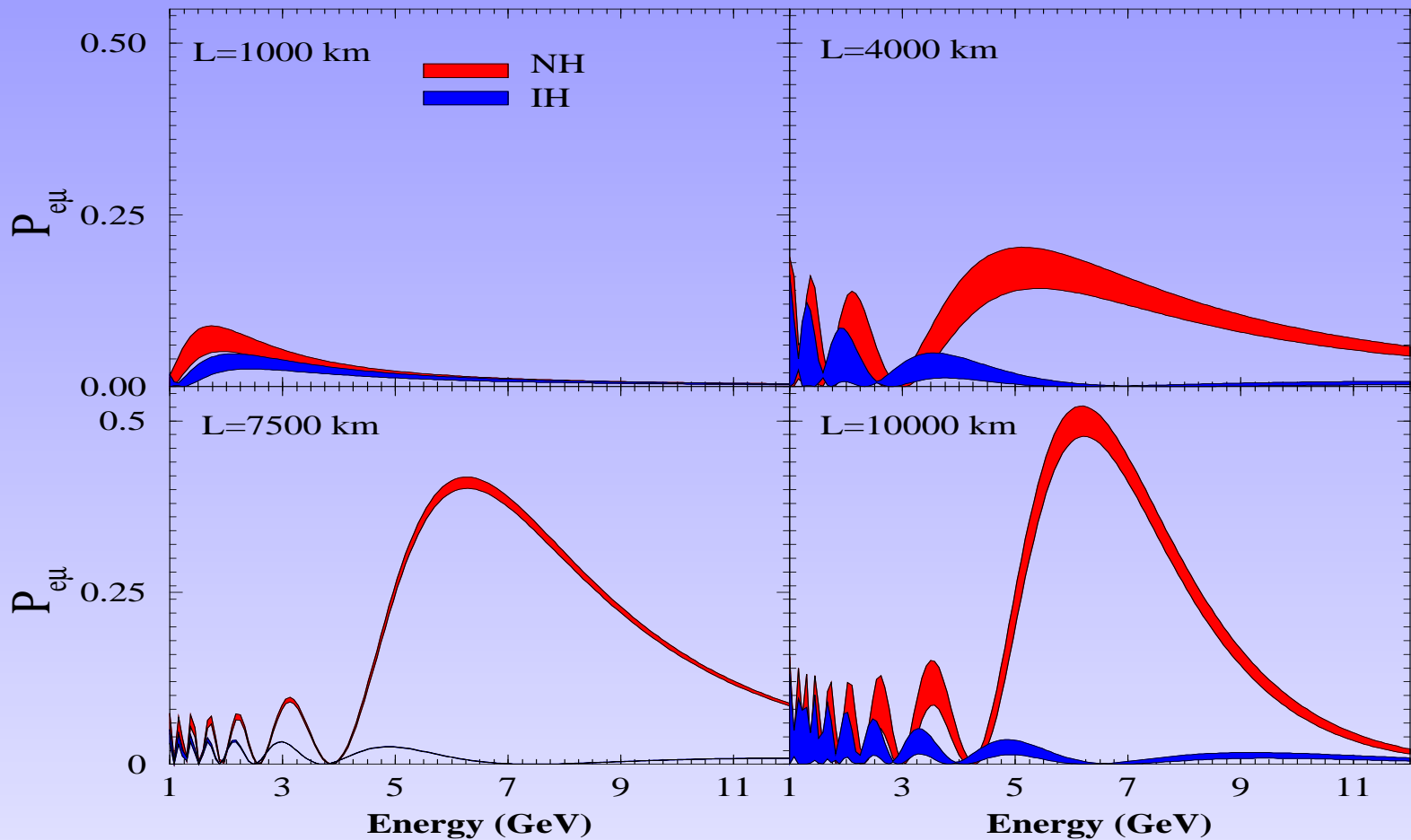
$$\sin^2 \theta_{12} = 0.31$$

$$\delta_{CP} = 0$$

Chosen benchmark values of oscillation parameters,
except $\sin^2 2\theta_{13}$

Exact 3-flav osc. prob using **PERM profile**

Transition Probability $P_{e\mu}$

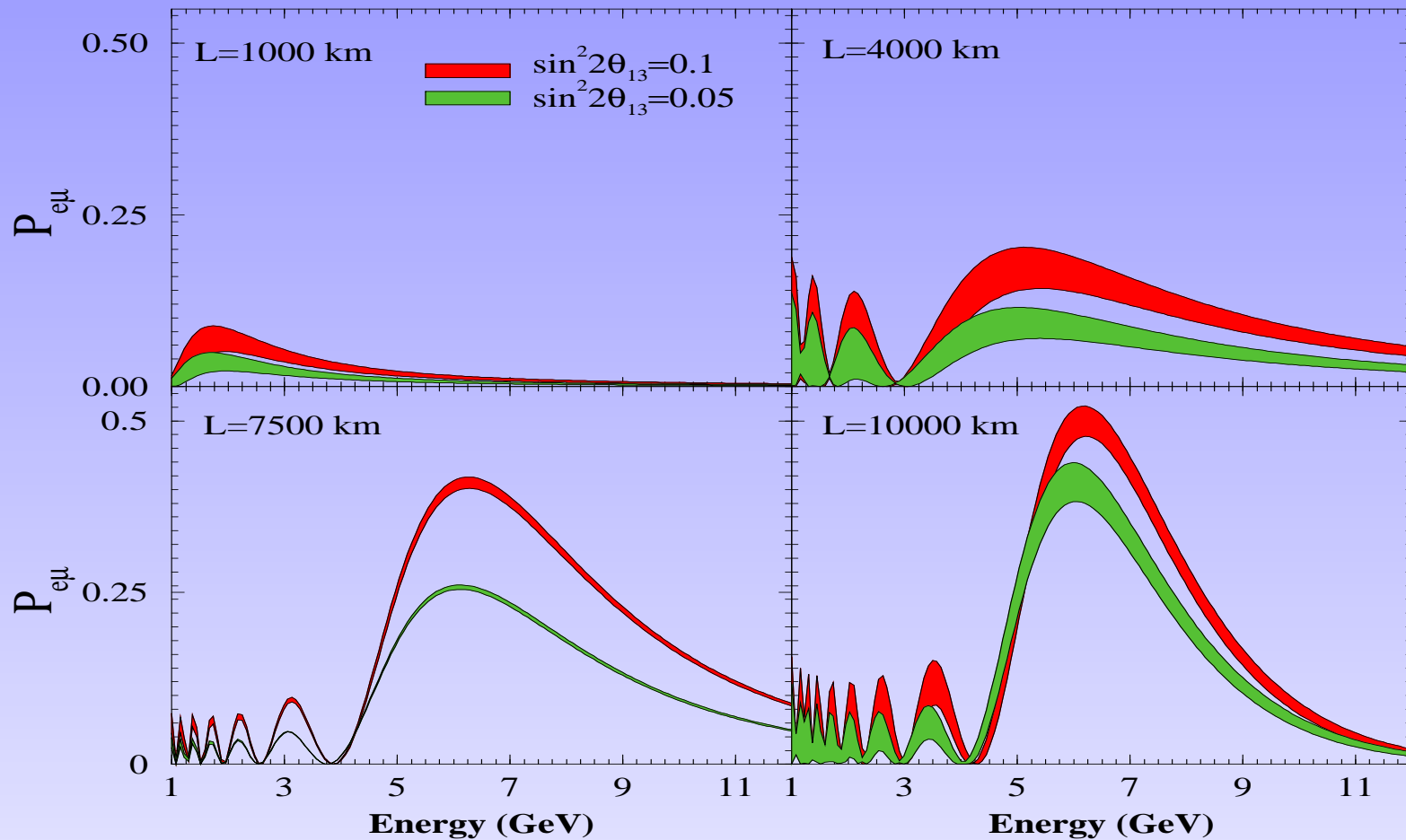


Agarwala, Choubey, Raychaudhuri, hep-ph/0610333

Normal .vs. Inverted hierarchy

$\sin^2 2\theta_{13} = 0.1$

Transition Probability $P_{e\mu}$



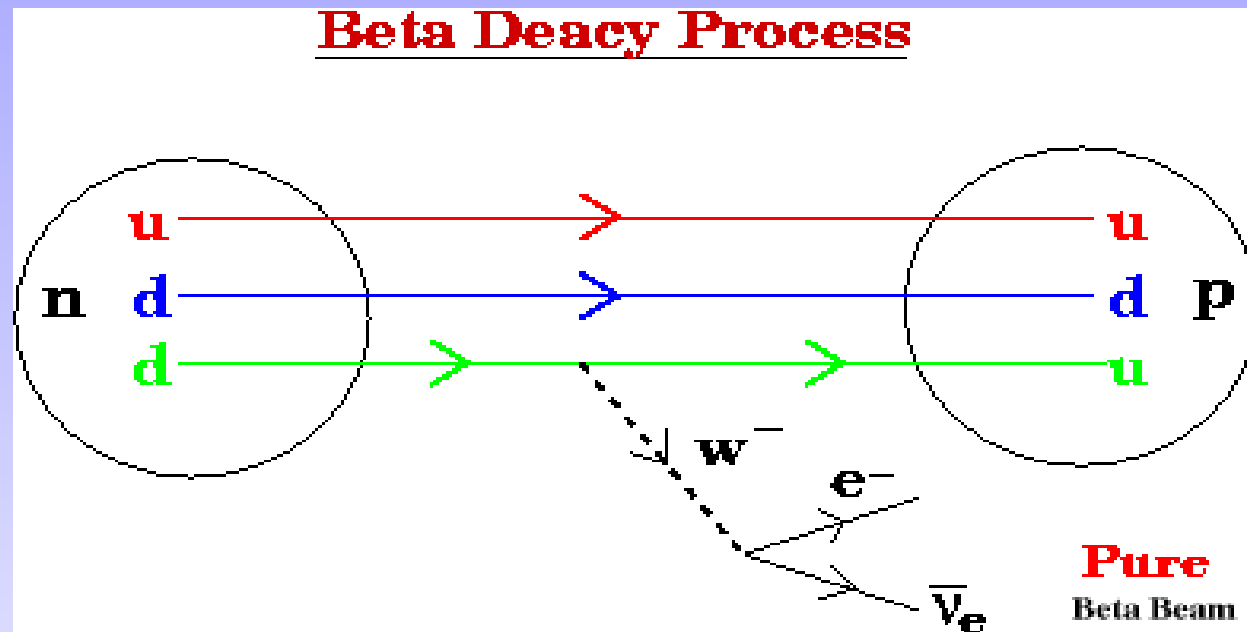
Agarwala, Choubey, Raychaudhuri, hep-ph/0610333

Two different values of $\sin^2 2\theta_{13}$

Normal hierarchy

What is Beta Beam?

A pure, intense, collimated beam of ν_e or $\bar{\nu}_e$,
essentially background free



P. Zucchelli, Phys. Lett. B 532 (2002) 166

Beta decay of completely ionized, radioactive ions
circulating in a storage ring. No contamination of
other types of neutrinos

Some positive features

- Known energy spectrum
- High intensity and low systematic errors
- High Lorentz boost of the parent ions \Rightarrow better collimation and higher energy of beam

Some positive features

- Known energy spectrum
- High intensity and low systematic errors
- High Lorentz boost of the parent ions \Rightarrow better collimation and higher energy of beam
- Can be produced with existing CERN facilities or planned upgrades
- It can be operated simultaneously in the ν_e as well as $\bar{\nu}_e$ mode. The boost factors are fixed by the e/m ratio of the respective ions

Beta Beam : Ion sources

Ion	τ (s)	E_0 (MeV)	f	Decay fraction	Beam
$^{18}_{10}\text{Ne}$	2.41	3.92	820.37	92.1%	ν_e
^6_2He	1.17	4.02	934.53	100%	$\bar{\nu}_e$
^8_5B	1.11	14.43	600684.26	100%	ν_e
^8_3Li	1.20	13.47	425355.16	100%	$\bar{\nu}_e$

Comparison of different source ions

Low- γ design, useful decays in case of anti-neutrinos
can be 2.9×10^{18} /year and for neutrinos
 1.1×10^{18} /year

Larger total end-point energy, E_0 is preferred

Schematic Lay-out

• Baseline Beta Beam facility comprises these sections

– Proton Driver

- SPL (≈ 4 GeV)

– ISOL Target

- spallation neutrons or direct protons

– Ion Source

- pulsed ECR

– Acceleration

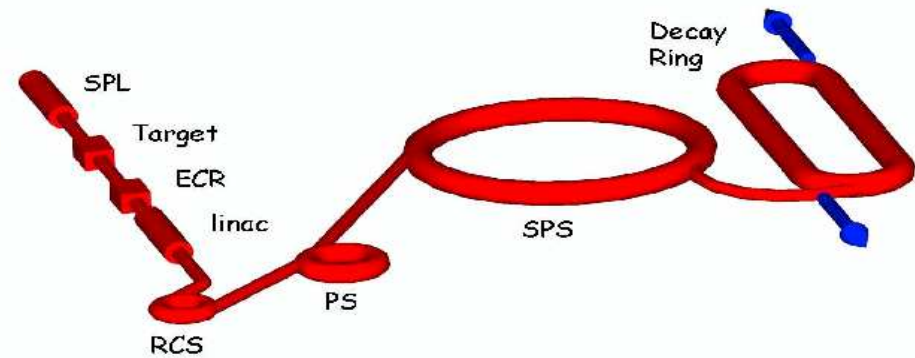
- linac, RCS, PS, SPS

– Decay Ring

- 7000 m; 2500 m straight

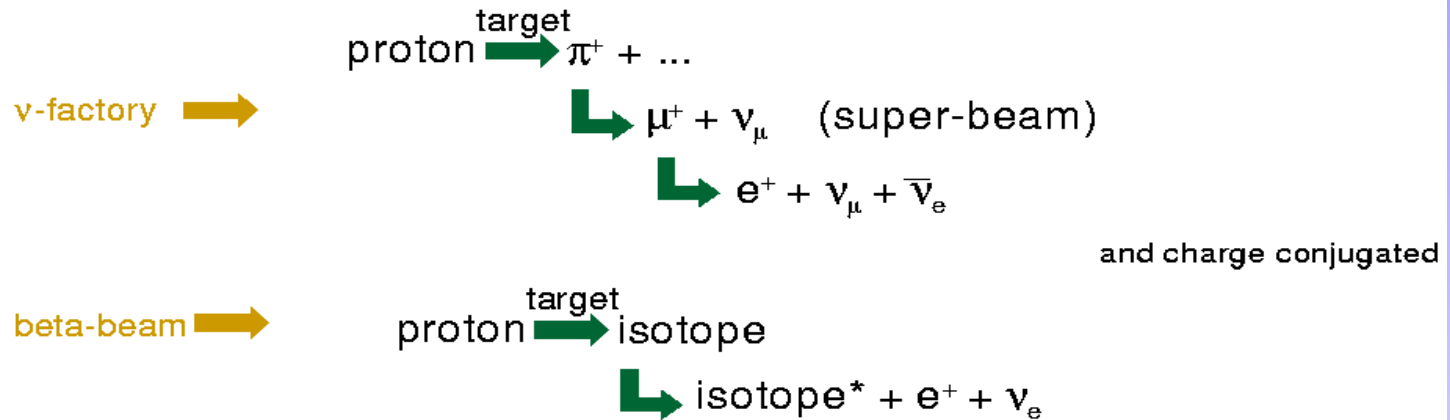
Baseline concept assumes
CERN PS, SPS

Use of Tevatron also
being considered

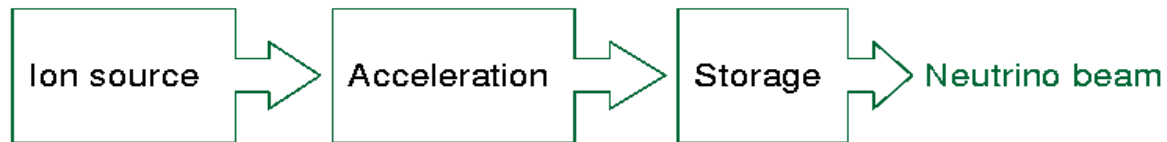


Proposed β -beam complex at CERN

Comparison with Nufact



v-factory uses beam of 4th generation.
Beta-beam uses 3rd generation beam.
Beta-beam is technically closer to existing/used accelerator technology.



Ultimate choice depends on future R&D

The India-based Neutrino Observatory

The INO/ICAL will be the world's first magnetized large mass iron calorimeter with interleaved Glass RPC detectors

Funding considerations in final stage

Location of INO

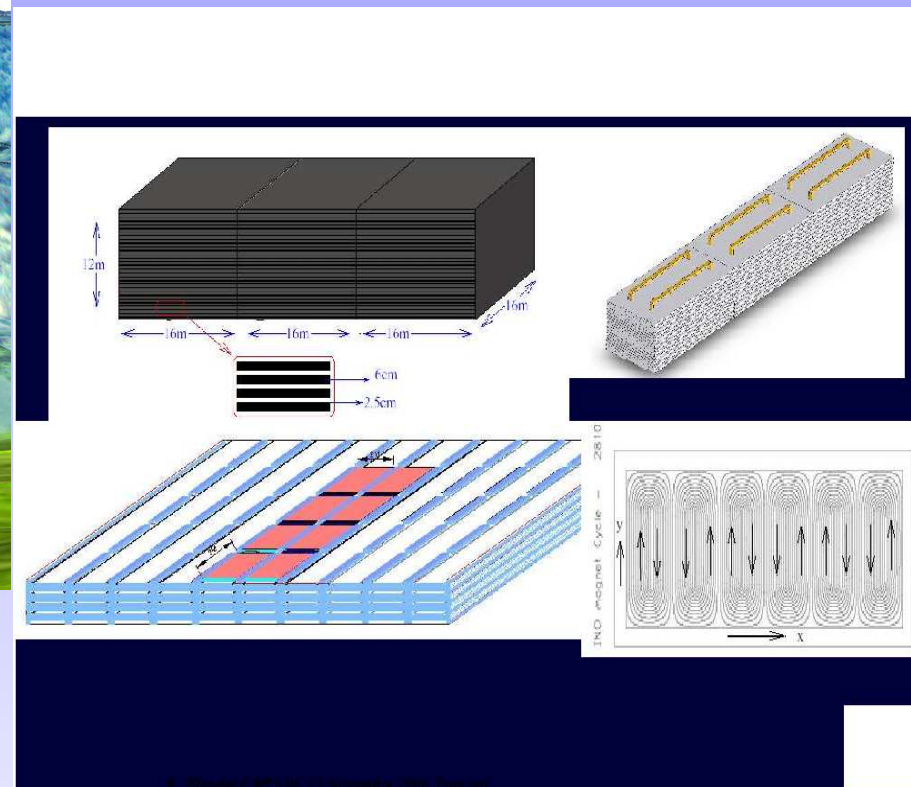
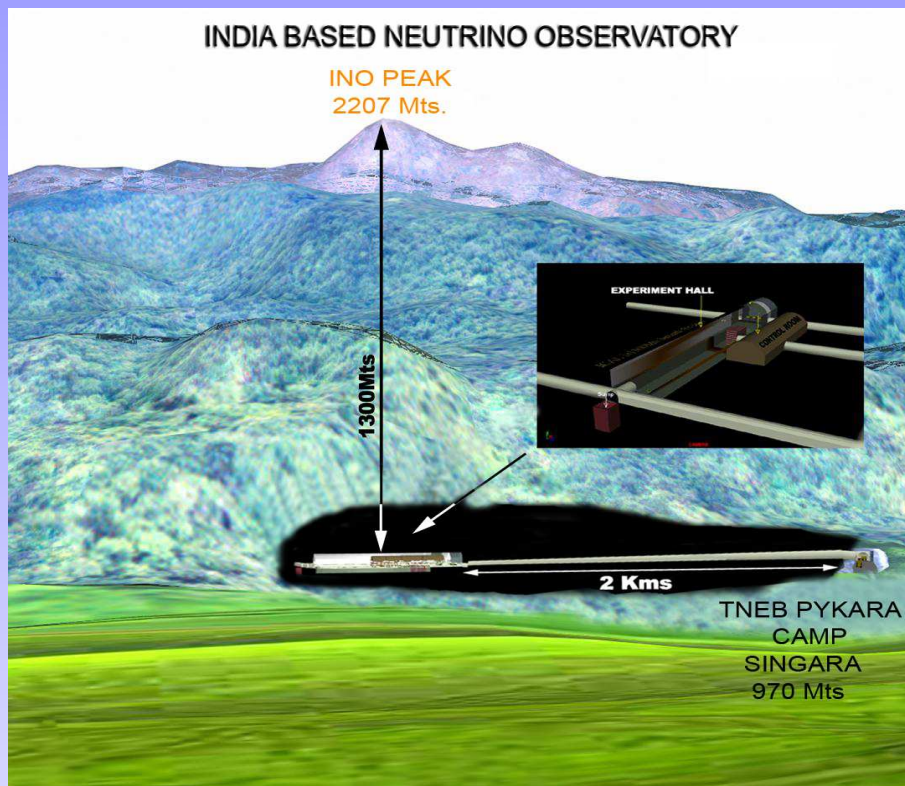


PUSHEP Site (Lat: N11.5°, Long: E76.6°)

PUSHEP-Bangalore: 250km

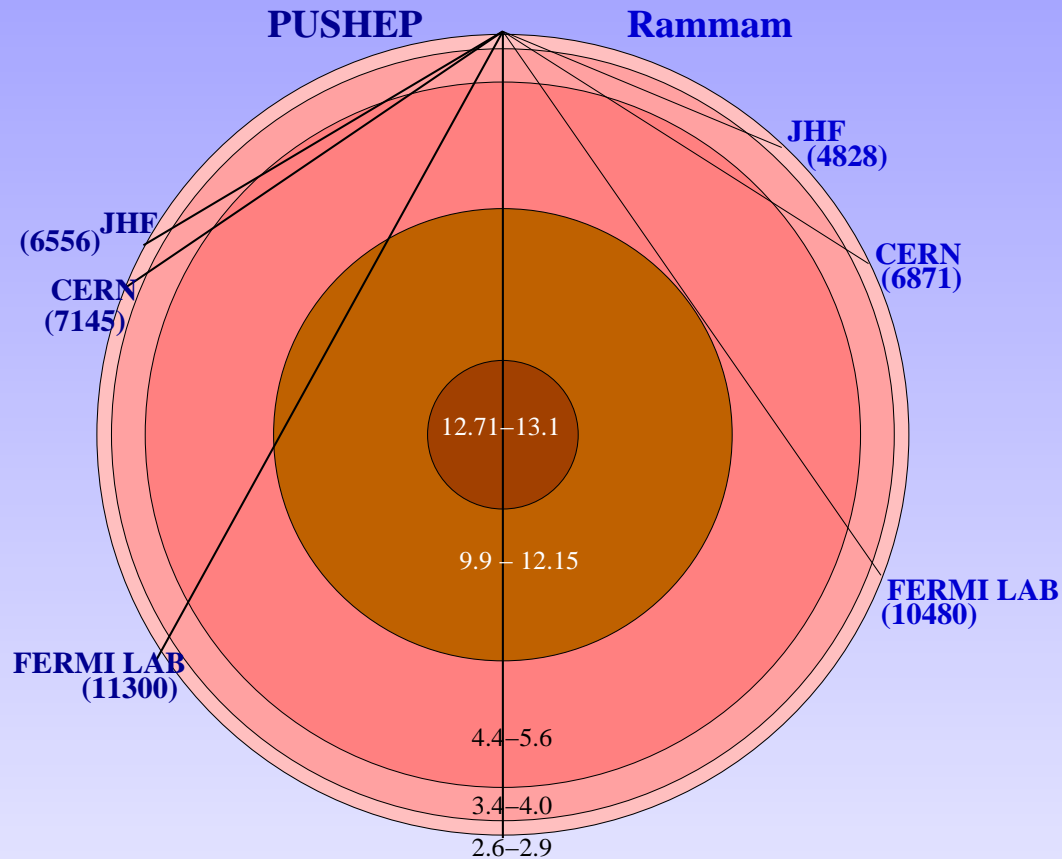
<http://www.imsc.res.in/~ino/>

ICAL Design



Spokesperson: Prof. N.K. Mondal, TIFR

INO : 2nd Phase



Artificial Source
Beta Beam ?
Neutrino Factory ?

$$L_{\text{magic}} = 7690 \text{ km}$$

- A magnetized Iron calorimeter (ICAL) detector with excellent efficiency of charge identification ($\sim 95\%$) and good energy determination
- Preferred location is Singara (PUSHEP) in the Nilgiris (near Bangalore), 7152 km from CERN
- A (50+50) Kton Iron detector
- Oscillation signal is the muon track ($\nu_e \rightarrow \nu_\mu$ channel)

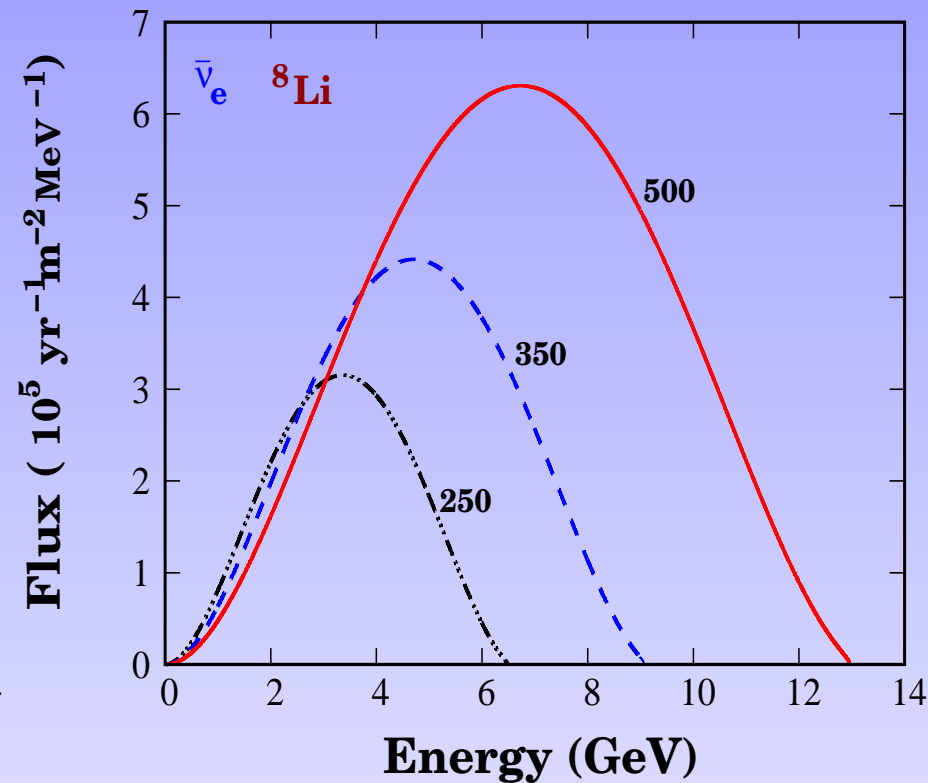
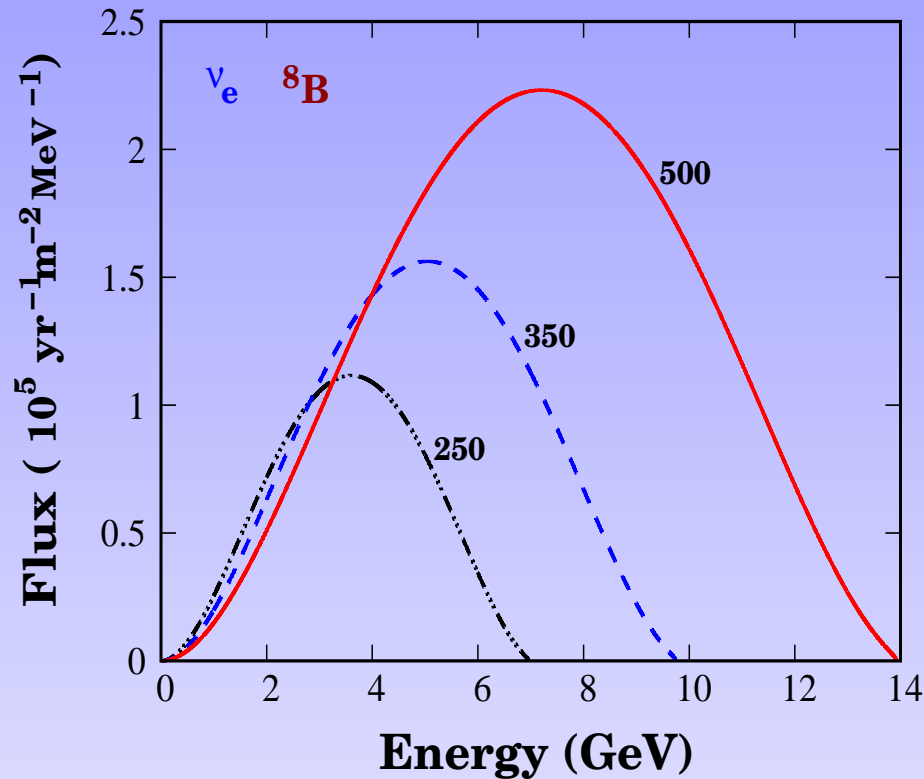
Detector assumptions

Total Mass	50 kton
Energy threshold	1.5 GeV
Detection Efficiency (ϵ)	60%
Charge Identification Efficiency (f_{ID})	95%

Detector characteristics used in the simulations

We assume a Gaussian energy resolution function
with $\sigma = 0.15E$

β -beam flux at INO-ICAL



Agarwalla, Choubey, Raychaudhuri, hep-ph/0610333

Boosted on-axis spectrum of ν_e and $\bar{\nu}_e$
at the far detector assuming no oscillation

CERN - INO Long Baseline

$$L_{\text{CERN-INO}} = 7152 \text{ km}$$

- The longer baseline captures a matter-induced contribution to the neutrino parameters, essential for probing the sign of Δm_{31}^2

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- The longer baseline captures a matter-induced contribution to the neutrino parameters, essential for probing the sign of Δm_{31}^2
- The CERN - INO baseline, close to the ‘Magic’ value, ensures essentially no dependence of the final results on δ_{CP} . This ‘Magic’ value is independent of E
- This permits a clean measurement of θ_{13} avoiding the degeneracy issues which plague other baselines

Resonance in matter effect

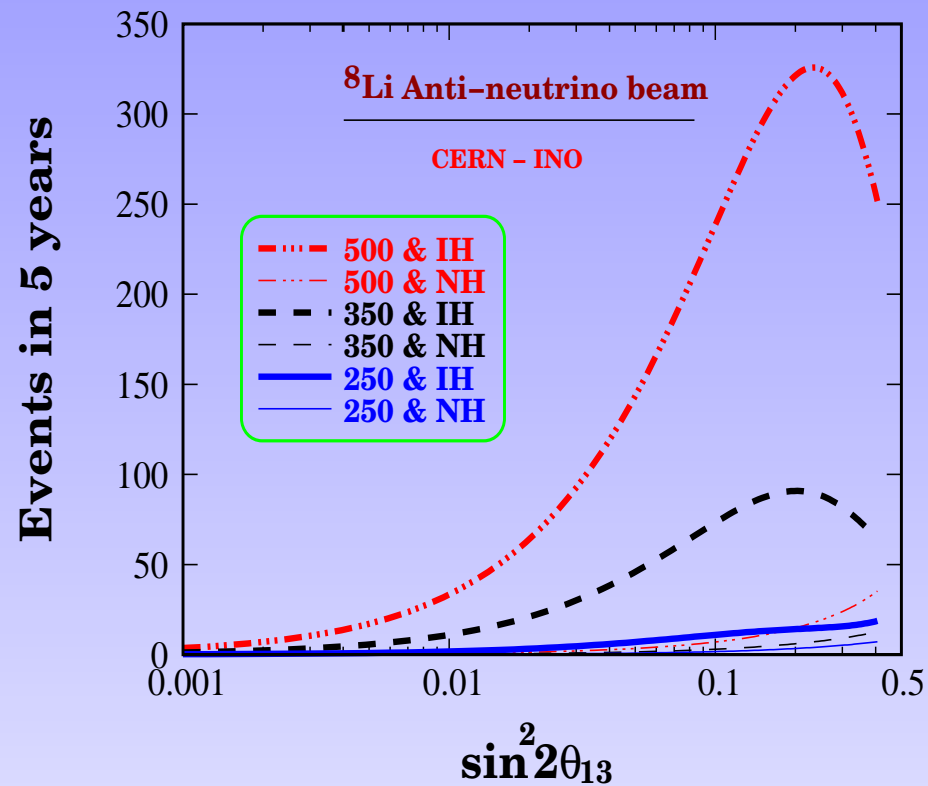
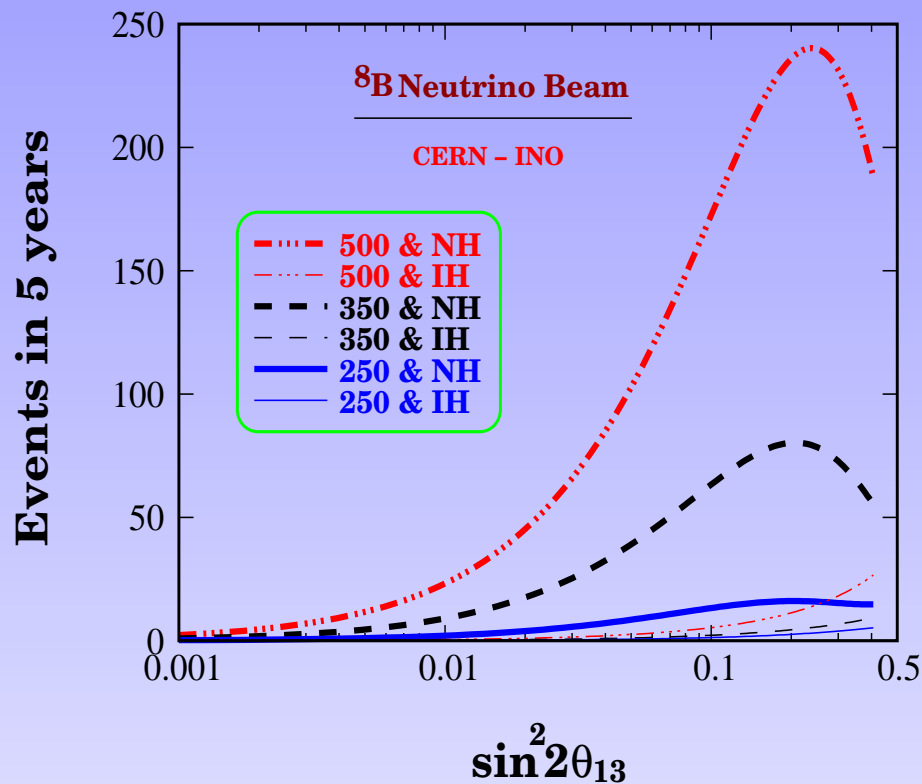
- The very long CERN - INO baseline provides an excellent avenue to pin-down matter induced contributions
- In particular, a resonance occurs at

$$E_{res} \equiv \frac{|\Delta m_{31}^2| \cos 2\theta_{13}}{2\sqrt{2}G_F N_e}$$

$$= 6.1 \text{ GeV}$$

with $|\Delta m_{31}^2| = 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{13} = 0.1$ and $\rho_{av} = 4.13 \text{ gm/cc}$ (PREM) for the baseline of 7152 km

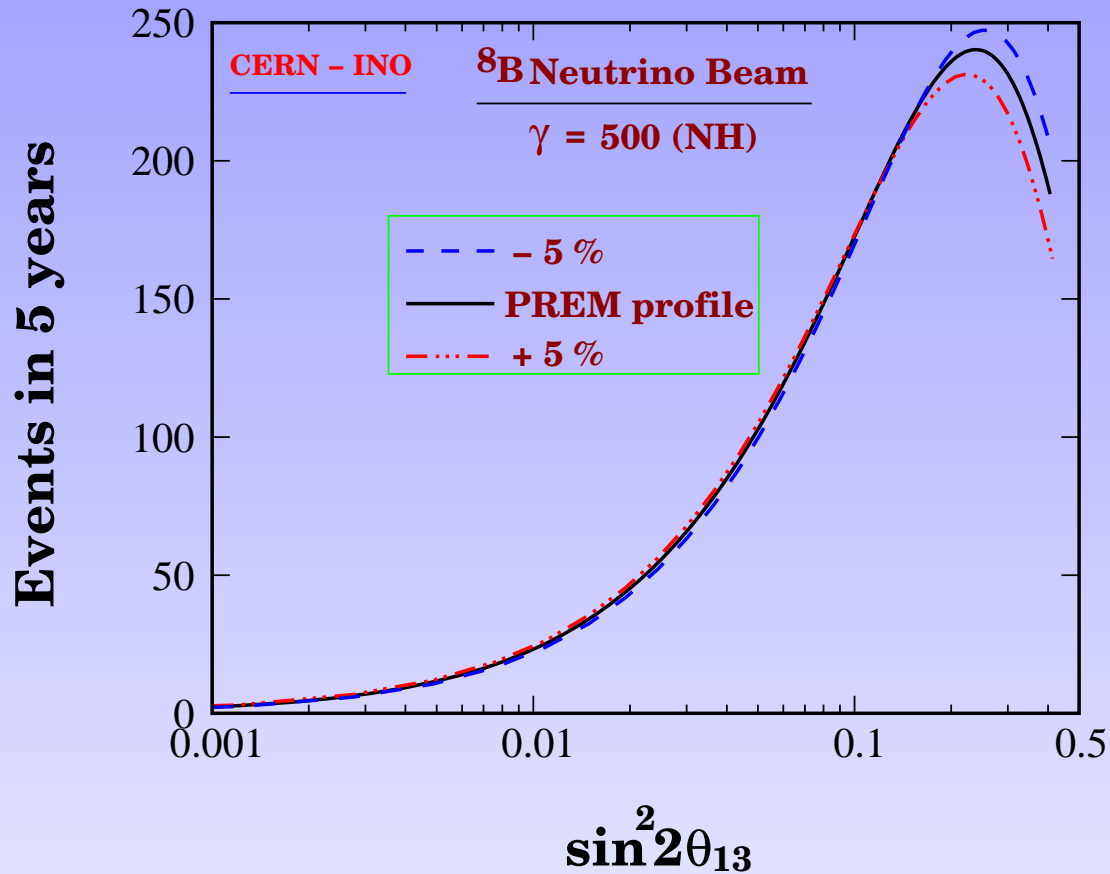
Event Rates in INO-ICAL



Agarwalla, Choubey, Raychaudhuri, hep-ph/0610333

Event rates sharply depend on mass ordering and θ_{13}

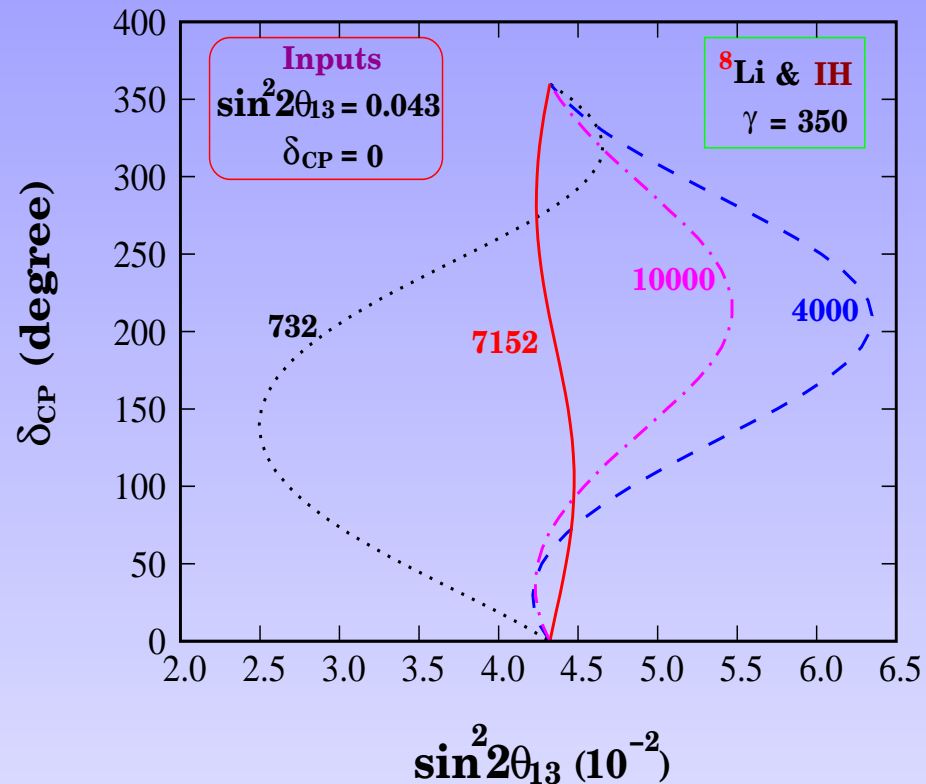
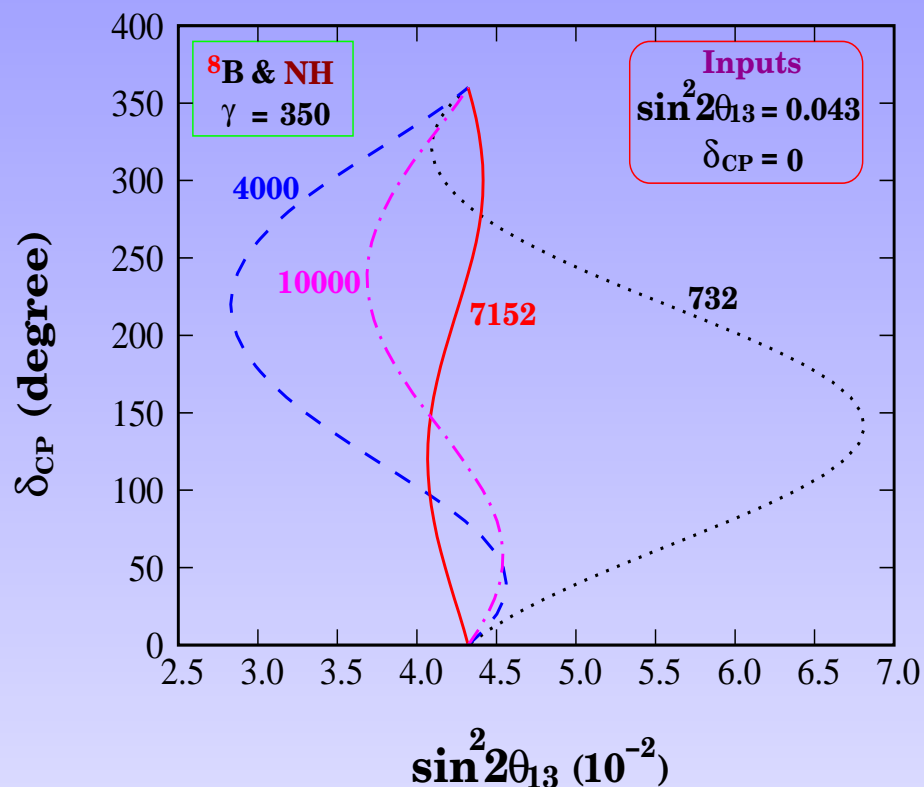
Event Rates (contd..)



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Sensitivity to matter profile

Iso-event curves



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At CERN-INO distance, the effect of δ_{CP} on the measurement of θ_{13} is less

The χ^2 function

Assume Poissonian distribution and define

$$\chi^2(\{\omega\}) = \min_{\xi_k} \left[2 \left(\tilde{N}^{th} - N^{ex} - N^{ex} \ln \frac{\tilde{N}^{th}}{N^{ex}} \right) + \sum_k \xi_k^2 \right]$$

$\{\omega\}$: oscillation parameters, $\{\xi_k\}$: “pulls”, where k runs over systematic uncertainties

$$\tilde{N}^{th}(\{\omega\}, \{\xi_k\}) = N^{th}(\{\omega\}) \left[1 + \sum_{k=1}^K \pi^k \xi_k \right] + \mathcal{O}(\xi_k^2)$$

Minimize χ^2 with respect to the pulls $\{\xi_k\}$ and finally marginalize over oscillation parameters by minimizing $\chi_{total}^2 = \chi^2 + \chi_{prior}^2$

The systematic errors

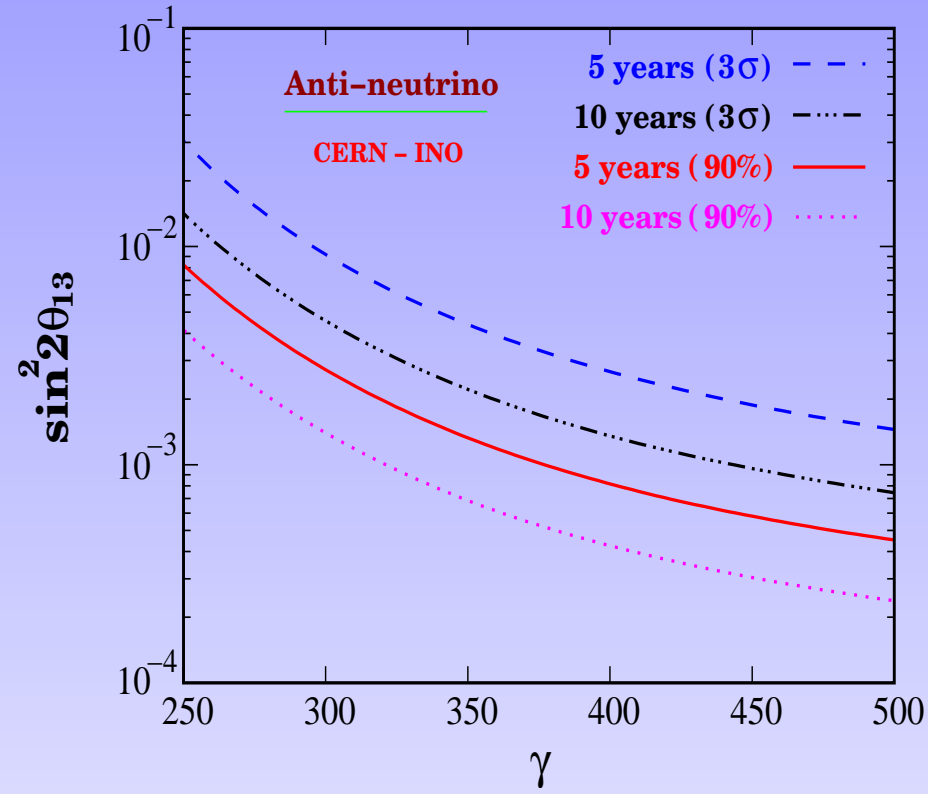
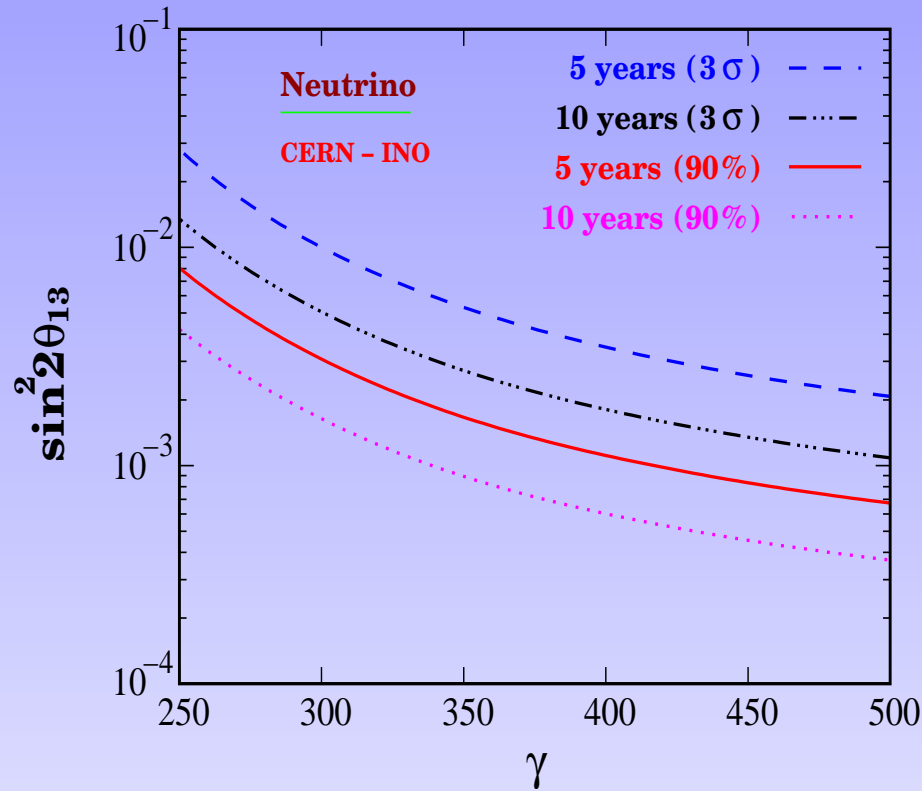
2% flux normalization error

10% error in the interaction cross section

Detector systematic uncertainty of 2%

No systematic error related to the shape of the energy spectrum. We work with the energy integrated total rates

Sensitivity to $\sin^2 2\theta_{13}$



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$\sin^2 2\theta_{13}$ limit below which experiment is insensitive

For ν_e ($\bar{\nu}_e$) true hierarchy is assumed normal (inverted)

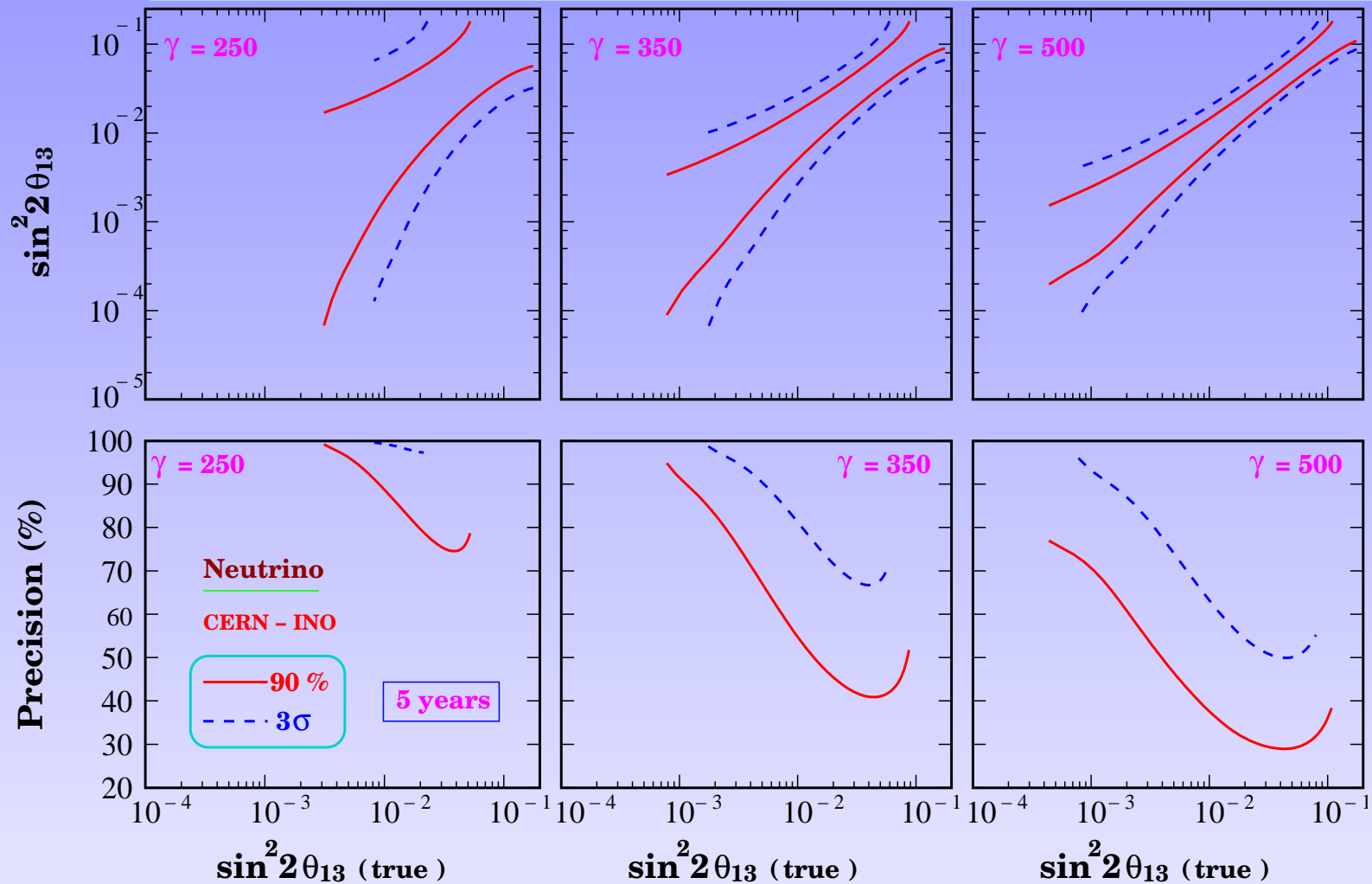
Marginalization over $|\Delta m_{31}^2|$, $\sin^2 2\theta_{23}$ and δ_{CP}

Upper limit on $\sin^2 2\theta_{13}$

At 3σ , $\sin^2 2\theta_{13} < 8.4 \times 10^{-4} (1.6 \times 10^{-3})$ with 80% detection efficiency and 10(5) years data in the neutrino mode assuming normal hierarchy as true with $\gamma = 500$

At 3σ , $\sin^2 2\theta_{13} < 1.1 \times 10^{-3} (2.1 \times 10^{-3})$ with 60% detection efficiency and 10(5) years data in the neutrino mode assuming normal hierarchy as true with $\gamma = 500$

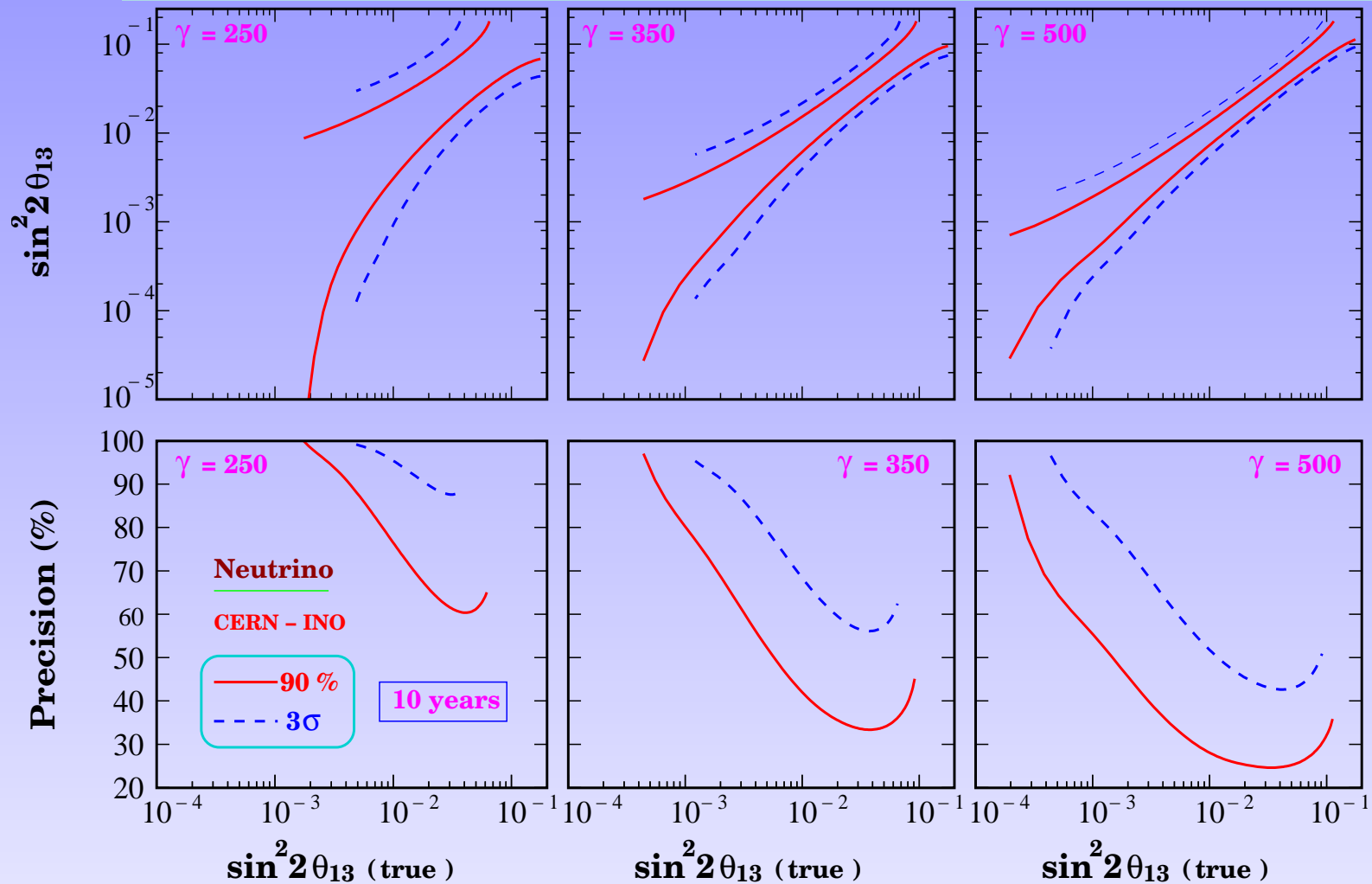
$\sin^2 2\theta_{13}$ precision (5 yrs)



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Upper panels show the band of “measured” values of $\sin^2 2\theta_{13}$ and lower ones depict the corresponding precision

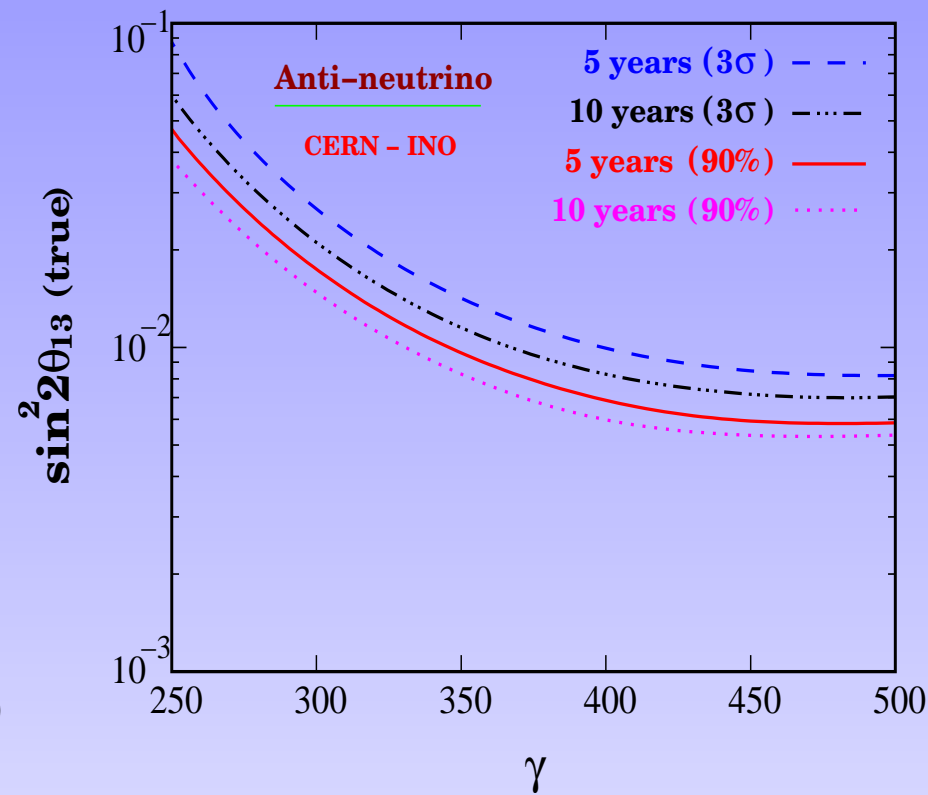
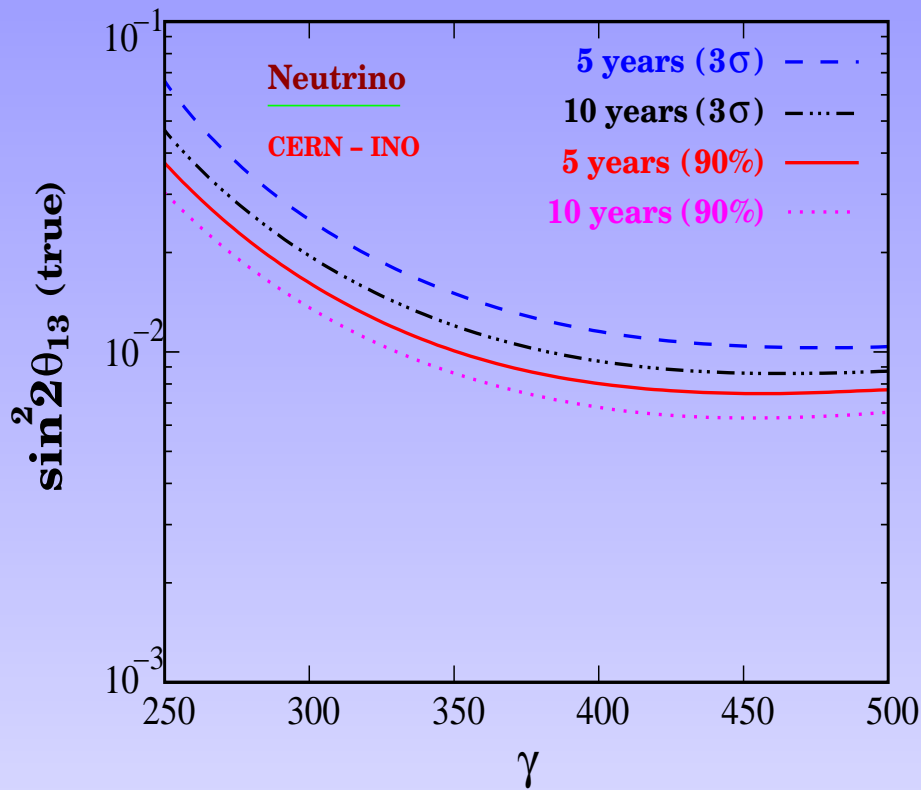
$\sin^2 2\theta_{13}$ precision (10 yrs)



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Upper panels show the band of “measured” values of $\sin^2 2\theta_{13}$ and lower ones depict the corresponding precision

Sensitivity to $Sgn(\Delta m_{31}^2)$



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Min value of $\sin^2 2\theta_{13}$ (true) as a function of γ to rule out wrong hierarchy. True hierarchy is normal (inverted) for ν_e ($\bar{\nu}_e$).

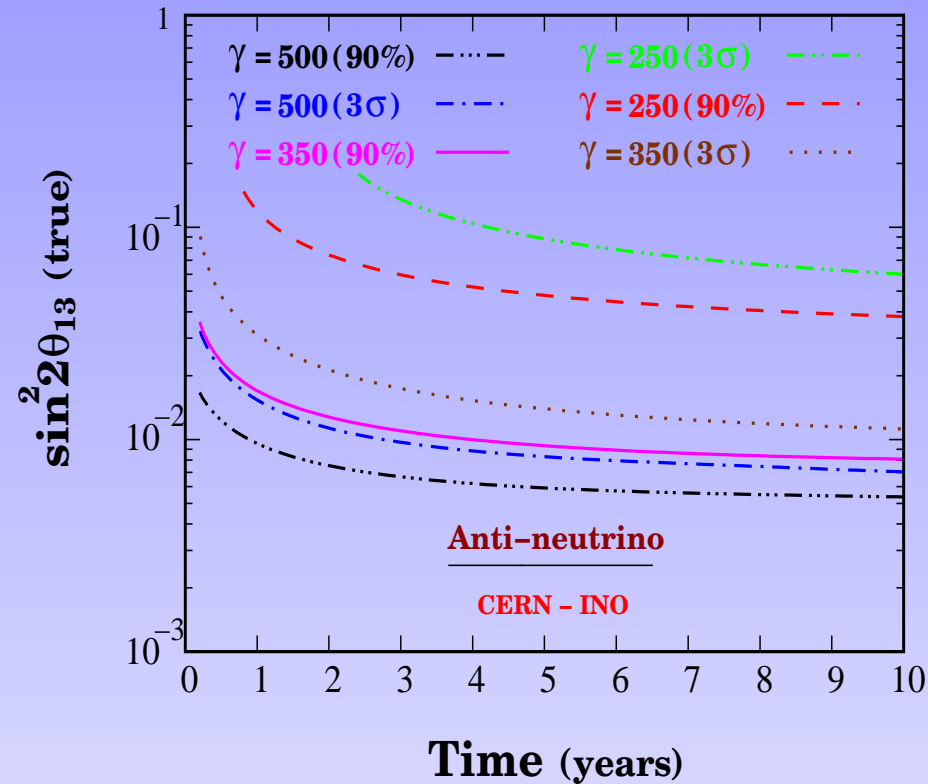
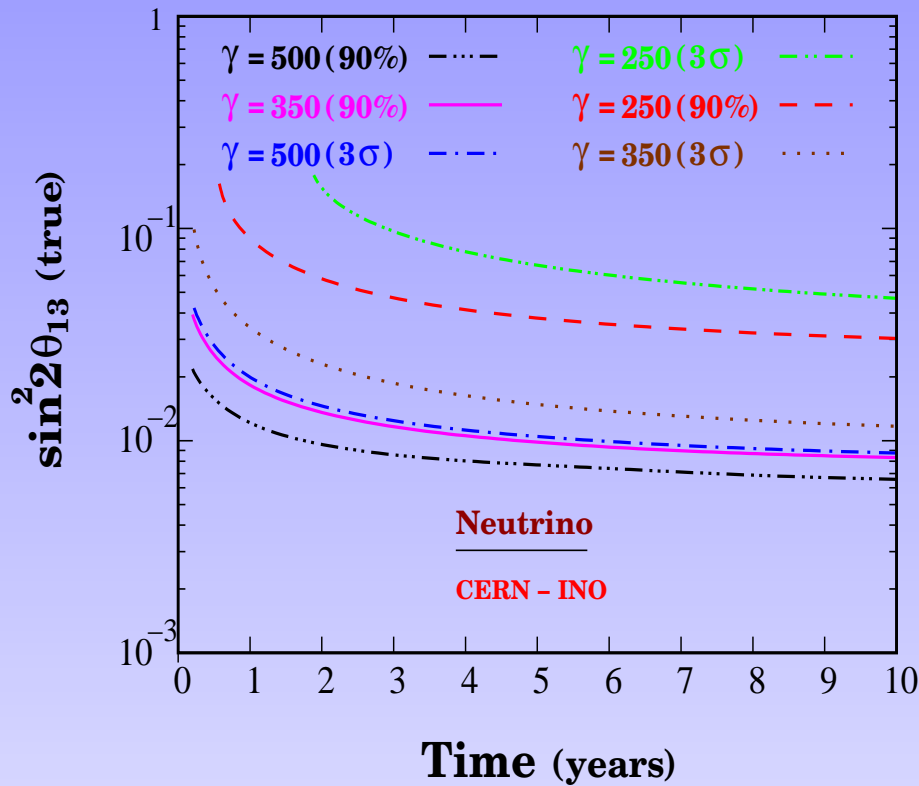
Marginalized over $|\Delta m_{31}^2|$, $\sin^2 2\theta_{23}$, δ_{CP} and $\sin^2 2\theta_{13}$

Sensitivity to mass ordering

Minimum value of $\sin^2 2\theta_{13} \rightarrow 8.5 \times 10^{-3} (9.8 \times 10^{-3})$ for which one can rule out inverted hierarchy at 3σ C.L. with 10(5) years of neutrino run assuming normal hierarchy as true hierarchy with $\gamma = 500$ and 80% detection efficiency

Minimum value of $\sin^2 2\theta_{13} \rightarrow 8.7 \times 10^{-3} (1.0 \times 10^{-2})$ for which one can rule out inverted hierarchy at 3σ C.L. with 10(5) years of neutrino run assuming normal hierarchy as true hierarchy with $\gamma = 500$ and 60% detection efficiency

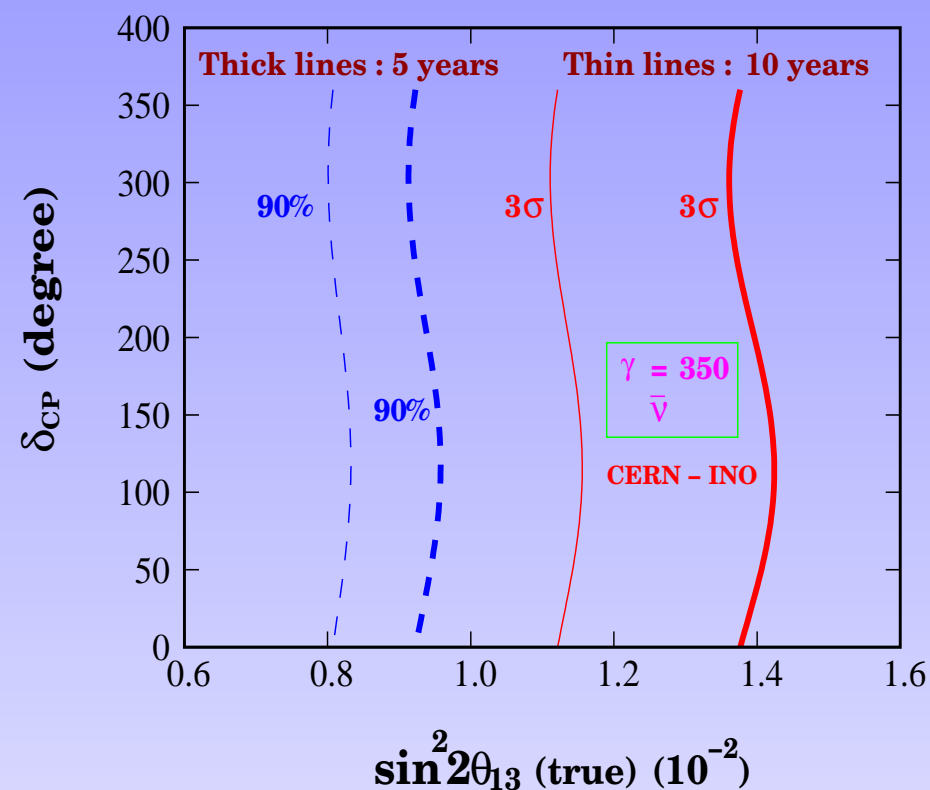
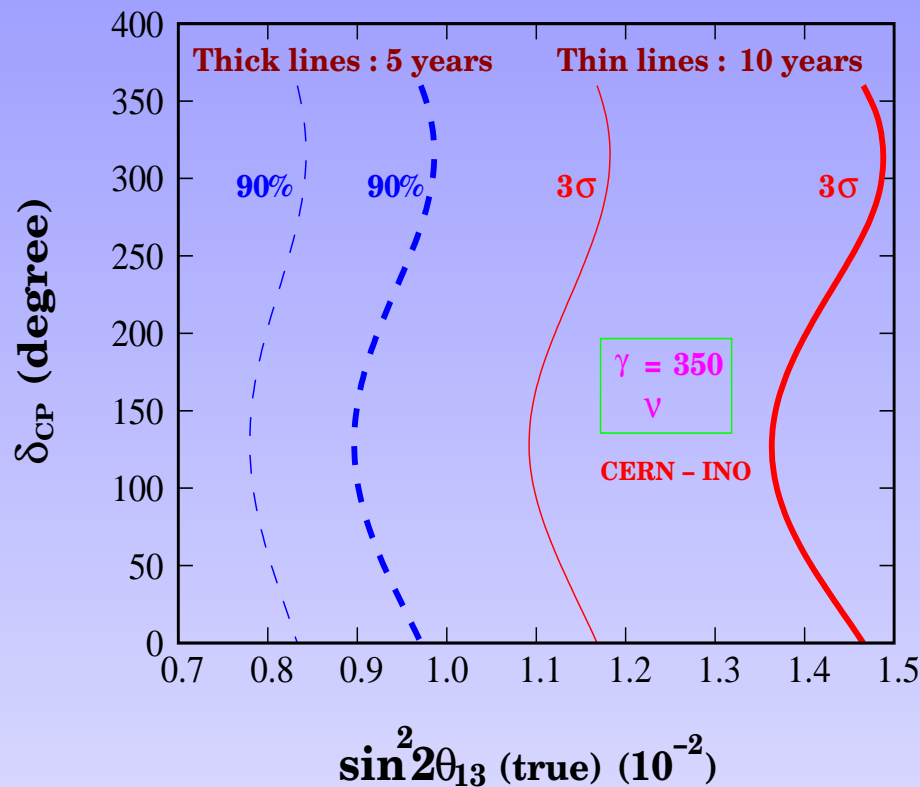
How much time we need?



Agarwalla, Choubey, Raychaudhuri, hep-ph/0610333

Just 9 months run in the ν_e mode to rule out the wrong inverted hierarchy at the 3σ C.L. with $\gamma = 500$ and 60% detection efficiency if $\sin^2 2\theta_{13}(\text{true}) = 5 \times 10^{-2}$ and the normal hierarchy is true

Impact of δ_{CP} on hierarchy



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The loss in hierarchy sensitivity due to the uncertainty in δ_{CP} is very marginal. Best sensitivity to hierarchy comes for $\delta_{CP} \simeq 125$ (300) in the ν ($\bar{\nu}$) mode

Conclusions

- **Ical@INO** → 50 Kt magnetized iron calorimeter
- **CERN-INO baseline (7152 Km)** - tantalizingly close to the magic baseline and hence will be free of the clone solutions

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- At this baseline we can get near-resonant matter effect for $E \approx 6 \text{ GeV} \Rightarrow$ possible with ${}^8_5\text{B}$ and ${}^8_3\text{Li}$
- Near-resonant matter effect gives largest possible $P_{e\mu} \Rightarrow$ enough statistics

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- Near-resonant matter effect gives largest possible $P_{e\mu} \Rightarrow$ enough statistics
- The CERN-INO Beta-Beam experiment is expected to give sensitivity to θ_{13} and $\text{Sgn}(\Delta m_{31}^2)$ better than all other rival proposals, apart from a high performance neutrino factory



Wait for Magic

!! Thank You !!