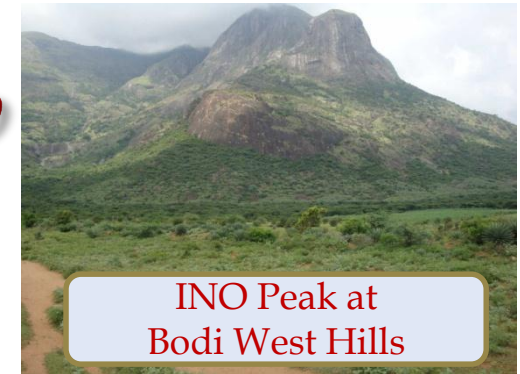
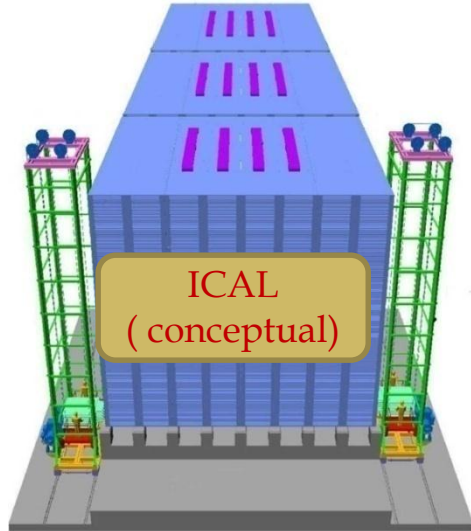


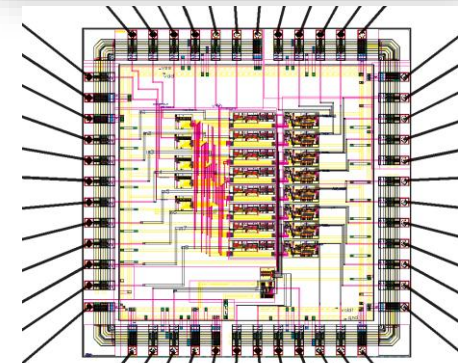
# Status of India-Based Neutrino Observatory (INO)

**B S Acharya,**  
**Tata Institute, Mumbai**  
*(on behalf of INO Collaboration)*



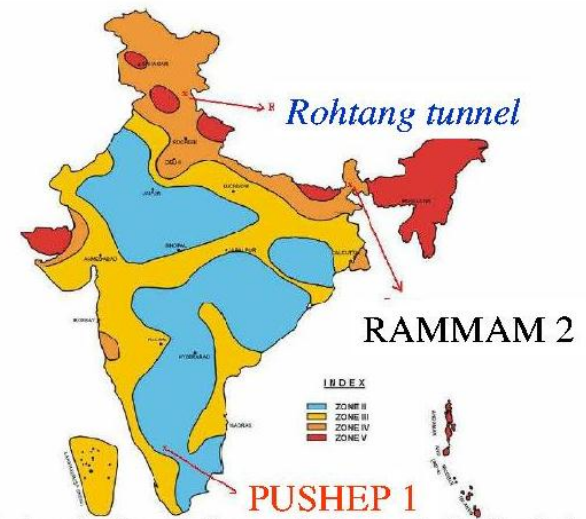
**Ahmedabad:** Physical Research Laboratory (PRL);  
**Aligarh:** Aligarh Muslim University (AMU);  
**Allahabad:** Harish Chandra Research Institute (HRI);  
**Bhuvanewar:** Institute of Physics (IOP); Utkal  
University; **Calicut:** University of Calicut;  
**Chandigarh:** Panjab University (PU); **Chennai:** Indian  
Institute of Technology, Madras (IITM); The Institute  
of Mathematical Sciences (IMSc); **Delhi:** Delhi  
University (DU); Jawaharlal Nehru University (JNU);  
**Kolkata:** Saha Institute of Nuclear Physics (SINP);  
University of Calcutta (CU); Variable Energy Cyclotron  
Centre (VECC); **Lucknow:** Lucknow University (LU);  
**Madurai:** American College; **Mumbai:** Bhabha  
Atomic Research Centre (BARC); Indian Institute of  
Technology, Bombay (IITB); Tata Institute of  
Fundamental Research (TIFR); **Mysuru:** University of  
Mysore (MU); **Srinagar:** University of Kashmir;  
**Varanasi:** Banaras Hindu University (BHU)

90+ Collaborators from  
22 Institutions



# INDIA

## States and Union Territories



Seismic zoning Map of India- issued by Bureau of Indian Standards, 2000

**BWH (9°58' N, 77°16' E)**  
 10 km from Theni (Railhead)  
 120 km from Madurai(Airport)

**Location of INO  
 (Bodi West Hills)**



# INO site at Pottipuram, Theni District

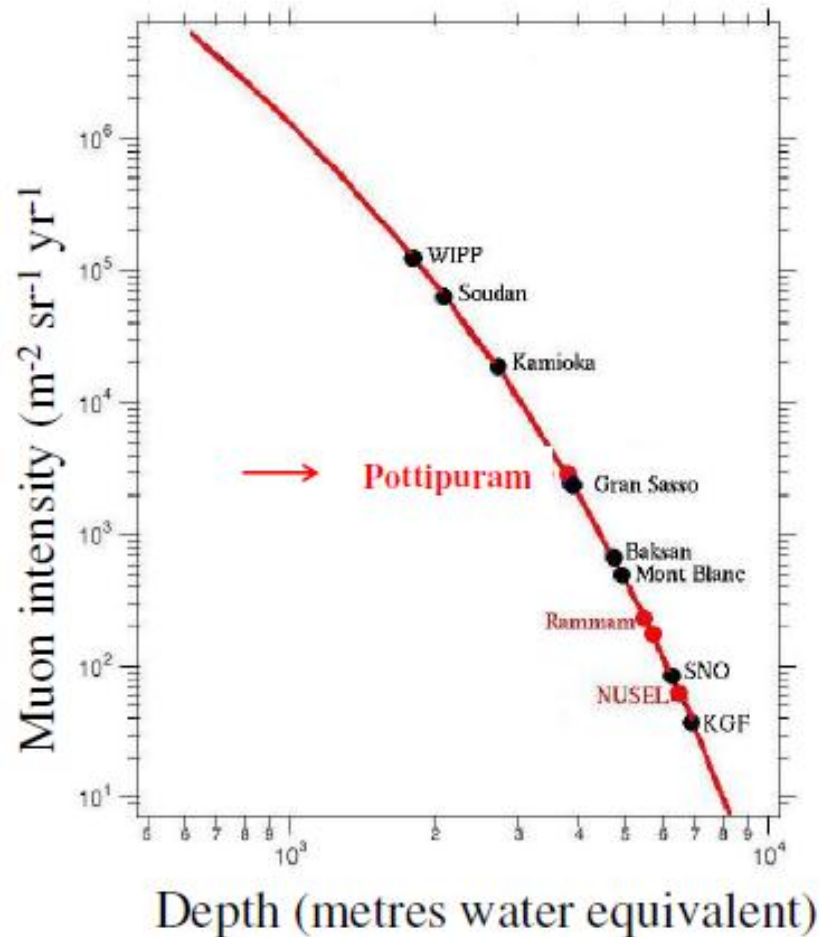


***Infrastructure at the site for water, electricity, approach road and local area development is going on.***



***~13 ha. of land at Madurai city to establish the Inter Institutional Centre for High Energy Physics ( IICHEP)***

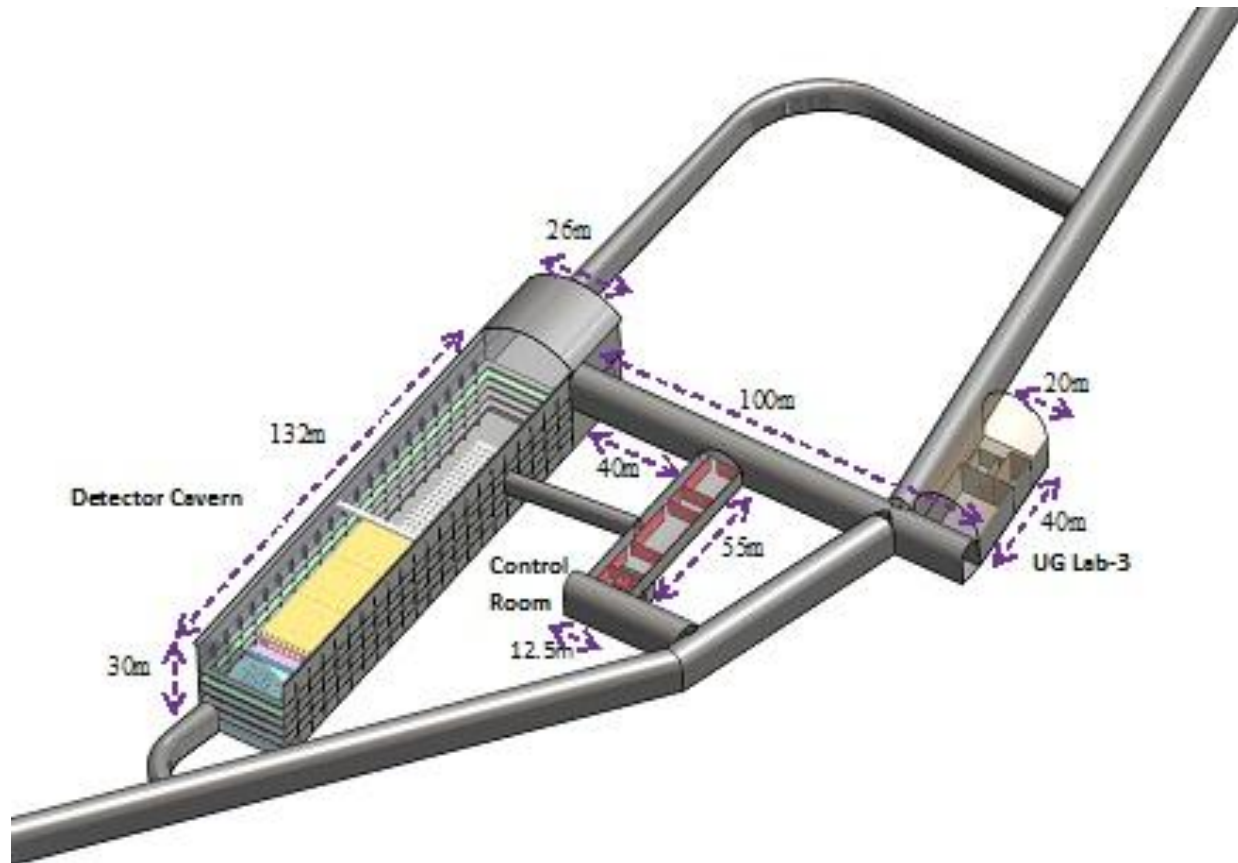
# The INO site at Pottipuram



- Cavern set in Charnockite rock under 1580 m peak – Bodi West Hills mountain
- Cavern access through 2m long adit
- 1289 m Vertical cover (1000 m all-round)
- Warm, low rain-fall area, low humidity through out the year, unusual wind speed in some seasons
- Flat terrain with good access to major roads

Close to magic baseline (~7500 Km)  
From CERN 7300 Km; JPARC 6500 Km;  
RAL 7600 Km;

# Underground Laboratory Layout



One large cavern for ICAL and 2 small cavern for other experiments such as Neutrino-less double beta decay (NDBD) and search for dark matter etc.

## Magnetized Iron Calorimeter (ICAL)

charge identification capability and good energy resolution

Excellent muon energy measurement, muon direction reconstruction and charge identification

Hadron shower reconstruction allows access to neutrino energy and high energy cosmic rays

### Motivation:

Atmospheric neutrinos provide wider range for E and L than any artificial neutrino sources

An ability to discriminate between neutrinos and anti-neutrinos enables efficient determination of neutrino mass ordering.

## ICAL - The physics goals

Determination of neutrino mass hierarchy

Accurate determination of atmospheric neutrino oscillation parameters

Reconfirm neutrino oscillations from distortion in L/E dependence,

Measure  $|\Delta m_{31}^2|$  and  $\text{Sin}^2 2\theta_{23}$

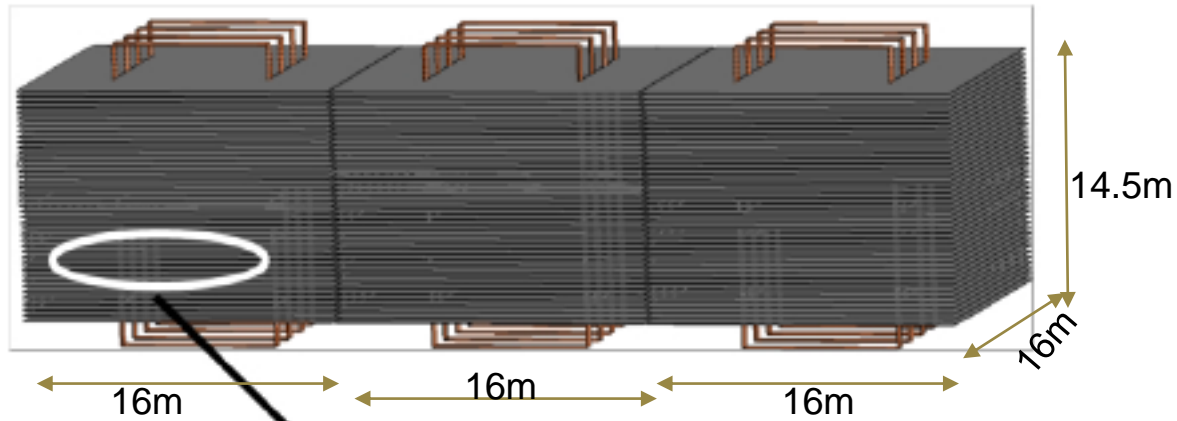
Deviation of  $\theta_{23}$  from maximality and octant of  $\theta_{23}$

Non-standard interactions, sterile neutrinos, CPT violations, LIV, long range forces, Ultra high energy muon fluxes etc.

Determination of CP violation in the lepton sector (with a future long baseline experiment with neutrino factory)

# INO-ICAL Detector Concept

151 LAYERS



50 k ton  
Magnetized  
Iron  
calorimeter

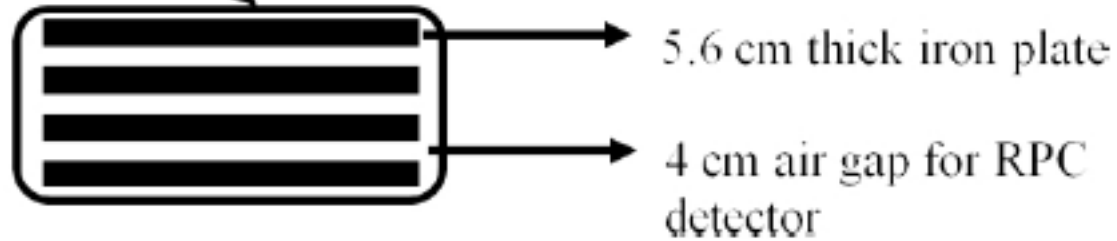
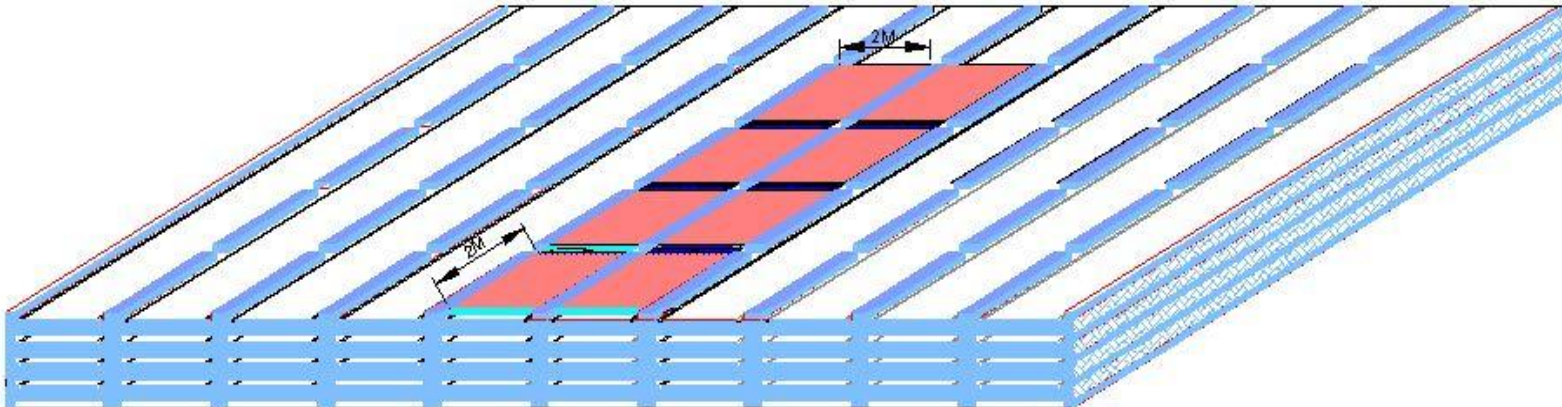


Figure 1.3: Schematic view of the 50 kt ICAL detector

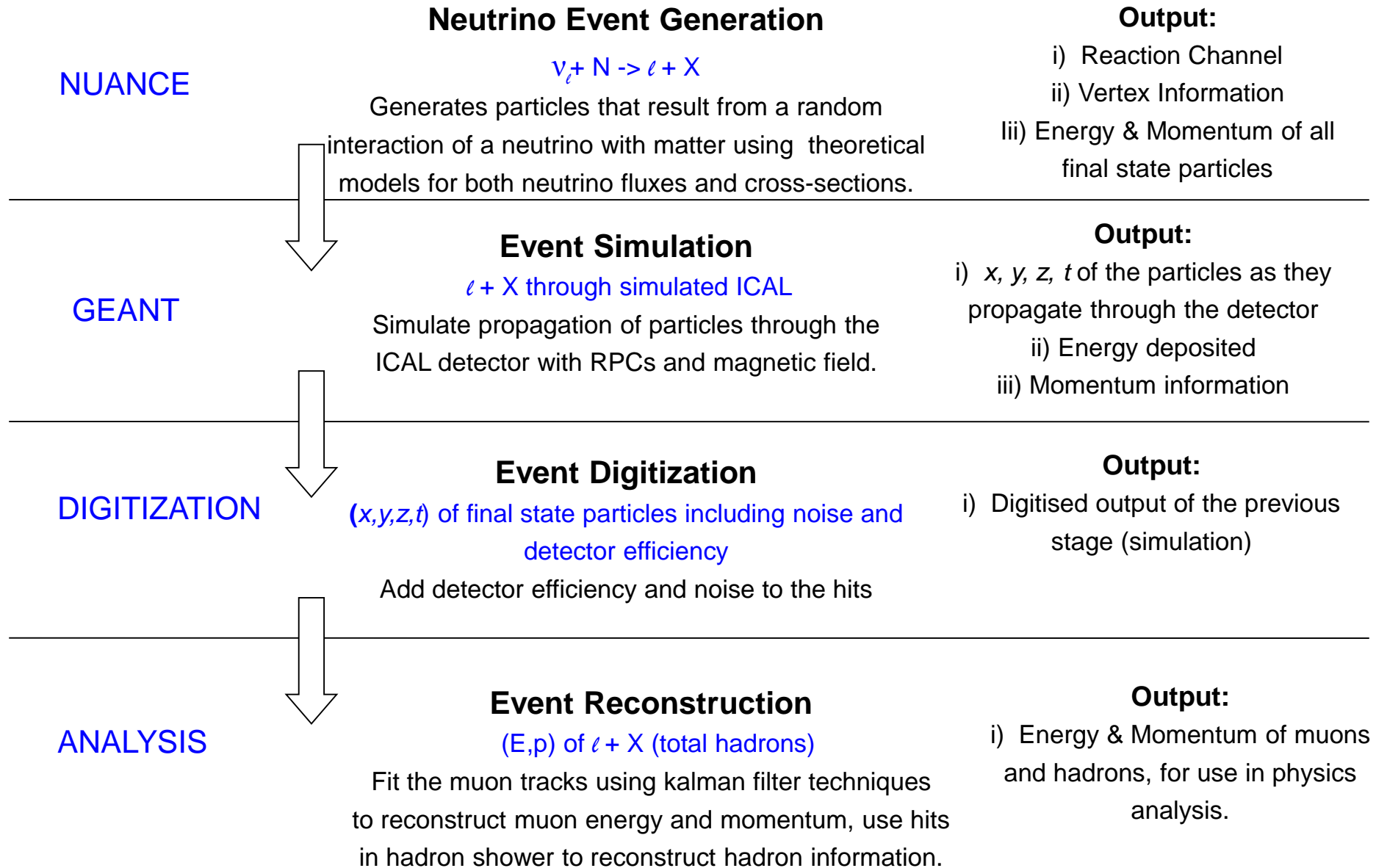




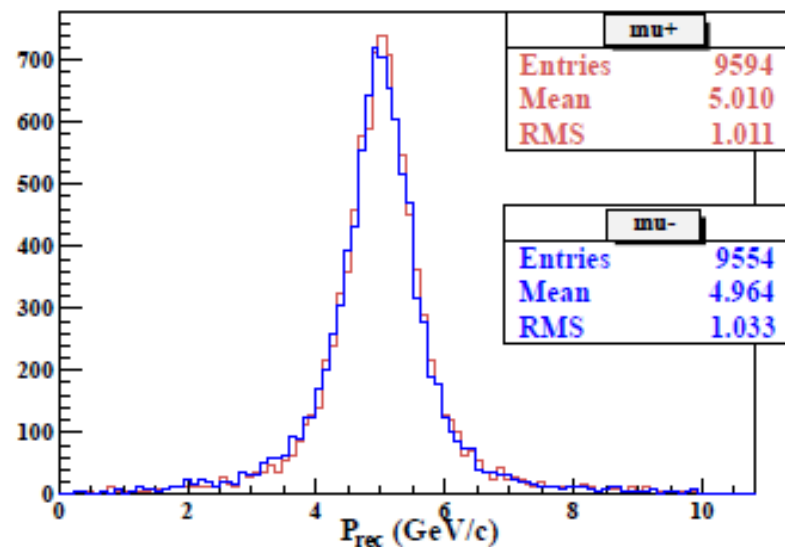
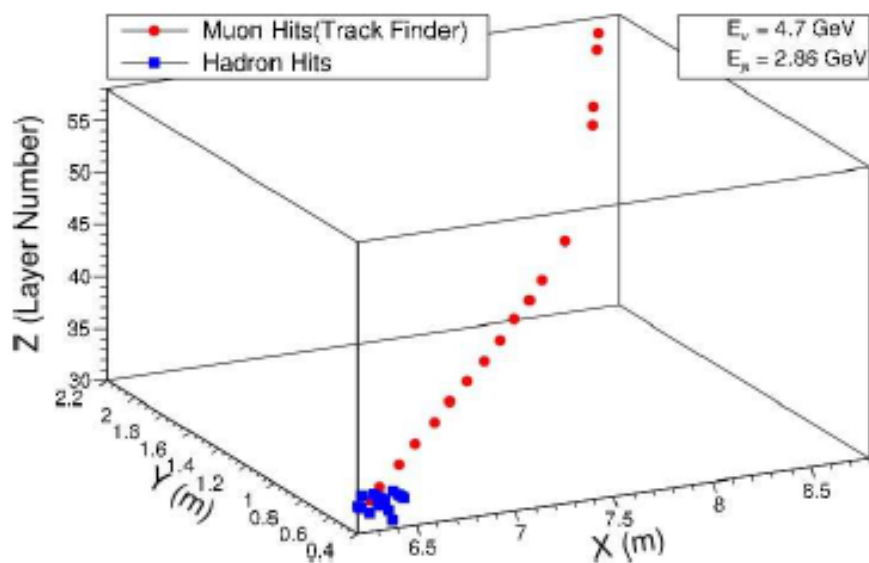
# INO-ICAL Detector Specifications

<i>No of modules</i>	<i>3 (50 k ton)</i>
<i>Module dimensions</i>	<i>16 m X 16 m X 14.5 m</i>
<i>Detector dimensions</i>	<i>48.4m X 16 m X 14.5 m</i>
<i>No of layers</i>	<i>151</i>
<i>Iron plate thickness</i>	<i>56 mm</i>
<i>Gap for RPC trays</i>	<i>40 mm</i>
<i>Magnetic field</i>	<i>1.5 Tesla</i>
<i>RPC unit dimensions</i>	<i>195 cm X 184 cm X 24 mm</i>
<i>Readout strip pitch</i>	<i>3 cm</i>
<i>No. of RPCs/Road/Layer</i>	<i>8</i>
<i>No. of Roads/Layer/Module</i>	<i>8</i>
<i>No. of RPC units/Layer</i>	<i>192</i>
<i>Total no of RPC units</i>	<i>28,992 (= 104,023 m<sup>2</sup>)</i>
<i>No of readout strips</i>	<i>3.710,976 (3.7 million)</i>

# Simulation Framework



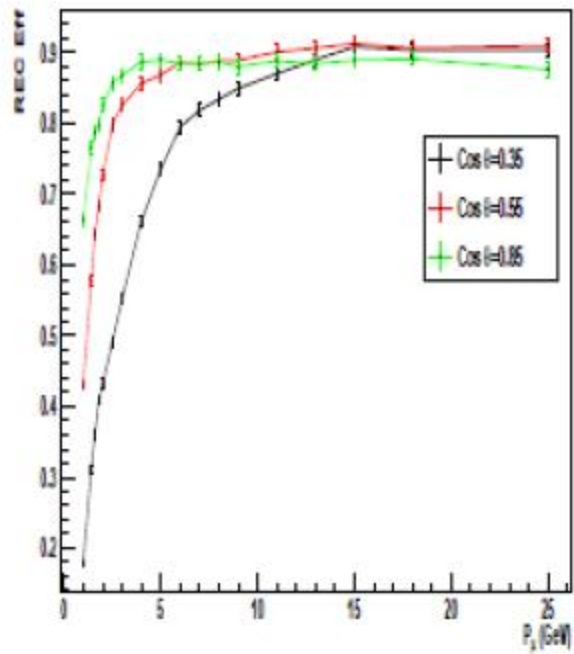
# ICAL Simulations



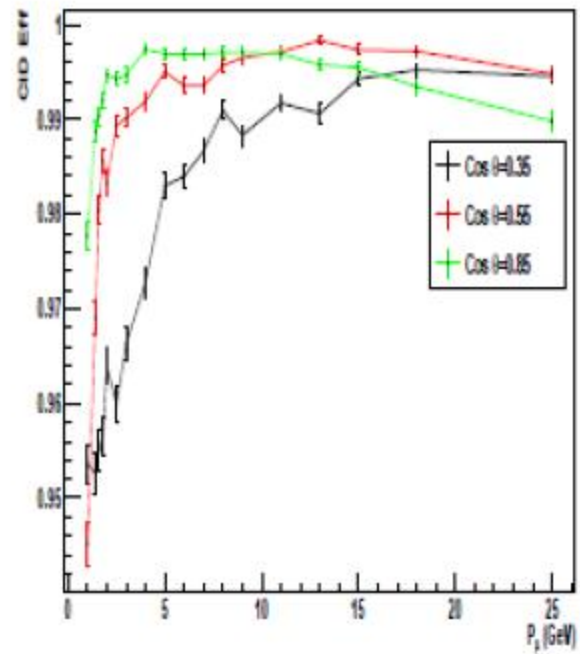
- GEANT4-based simulation of ICAL detector. Magnetic field map through simulations using MAGNET6.0 code.
- Neutrino events generated using NUANCE neutrino generator
- Muons leave long, clean tracks in detector. Calibrated through range or bending in magnetic field (Kalman filter)
- Hadrons are calibrated through the hits they leave: do not traverse many iron layers

# Detector performance: efficiencies

Reconstruction Efficiency

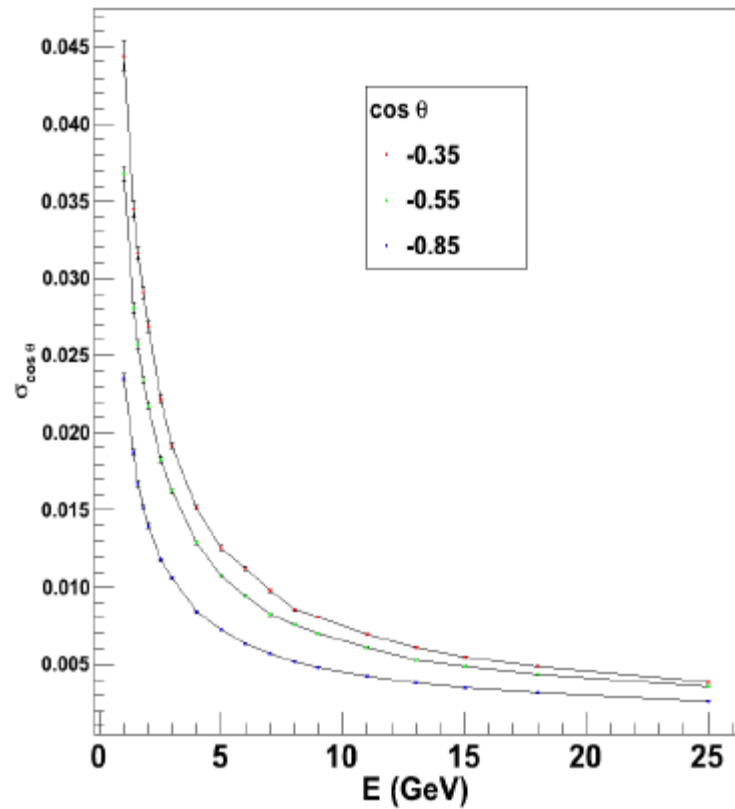
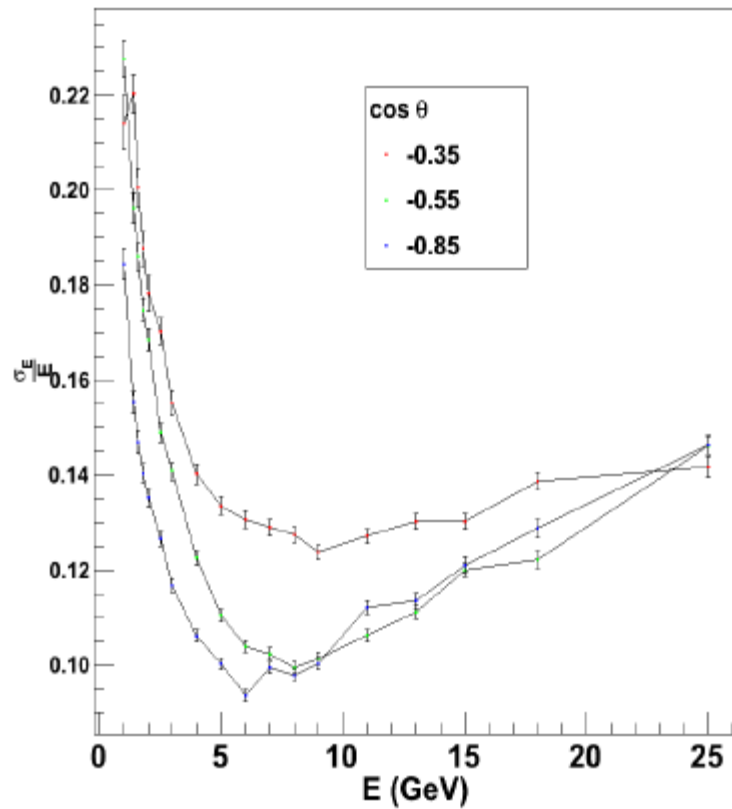


CID Efficiency



$\mu^-$

# Detector performance : resolutions



$\mu^-$

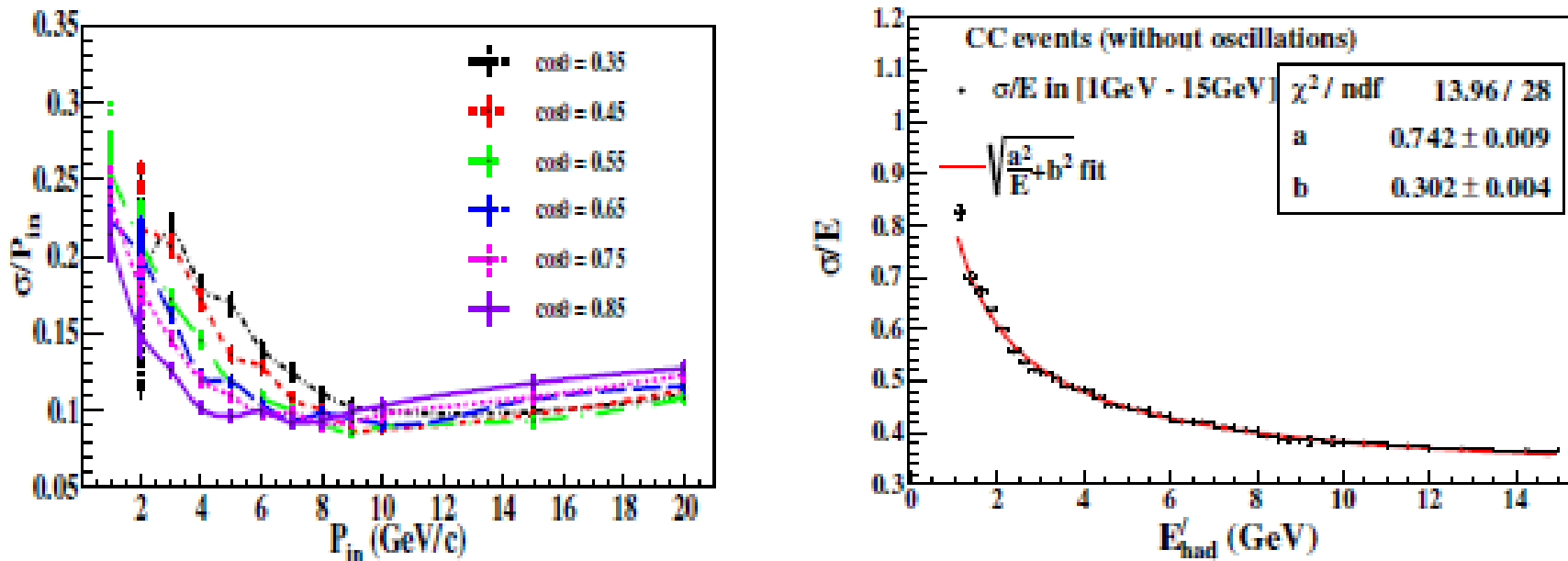


Figure 1: Left panel shows the momentum resolution of muons produced in the region  $0 < \phi < \pi/4$  (see Sec. 4.1), as functions of the muon momentum in different zenith angle bins [8]. Right panel shows the energy resolution of hadrons (see Sec. 4.2) as functions of  $E'_{had}$ , where events have been generated using NUANCE in different  $E'_{had}$  bins. The bin widths are indicated by horizontal error bars [9].

	$\mu^-$	$\mu^+$
Unoscillated	14311	5723
Oscillated	10531	4188
After Applying Reconstruction and CID Efficiencies	4941	2136
After Applying $(E, \cos\theta)$ Resolutions	5270	2278

Table 5.1: Number of muon events produced in CC  $\nu_\mu$  interactions at various stages of the analysis for an exposure of 50 kt  $\times$  10 years in the energy range 0.8–10.8 GeV.

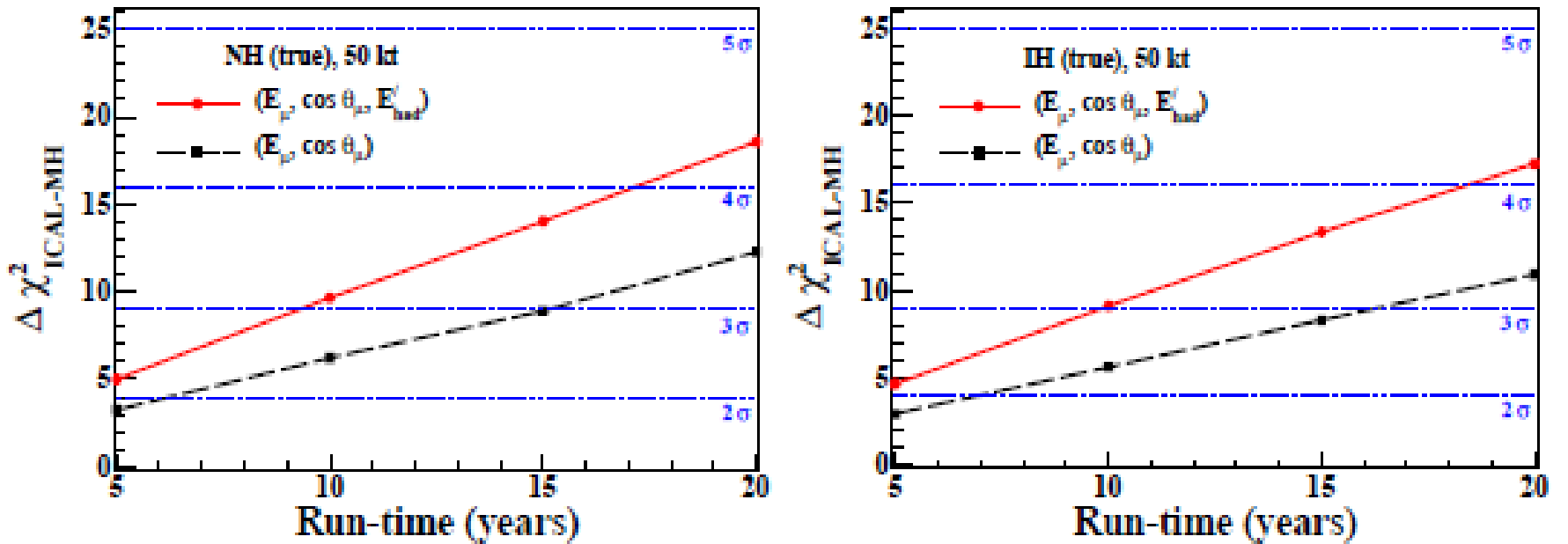


Figure 2: The hierarchy sensitivity of ICAL with input normal (left) and inverted (right) hierarchy including correlated hadron energy information, with  $|\Delta m_{\text{eff}}^2|$ ,  $\sin^2 \theta_{23}$  and  $\sin^2 2\theta_{13}$  marginalised over their  $3\sigma$  ranges [14]. Improvement with the inclusion of hadron energy is significant.

### Identifying mass hierarchy at all $\delta_{\text{CP}}$ values

The large range of path length of the atmospheric neutrinos makes ICAL insensitive to the CP phase  $\delta_{\text{CP}}$ , as a result its reach in distinguishing the hierarchy is also independent of the actual value of  $\delta_{\text{CP}}$  [18]. On the other hand the sensitivity of fixed-baseline experiments such as T2K and NO $\nu$ A is extremely limited if  $0 < \delta_{\text{CP}} < \pi$ . However adding of the ICAL information ensures that the hierarchy can be identified even in these unfavoured  $\delta_{\text{CP}}$  regions [11]. Of course in the  $\delta_{\text{CP}}$  regions favourable to the long baseline experiments, the ICAL data can only enhance the power of discriminating between the two hierarchies.



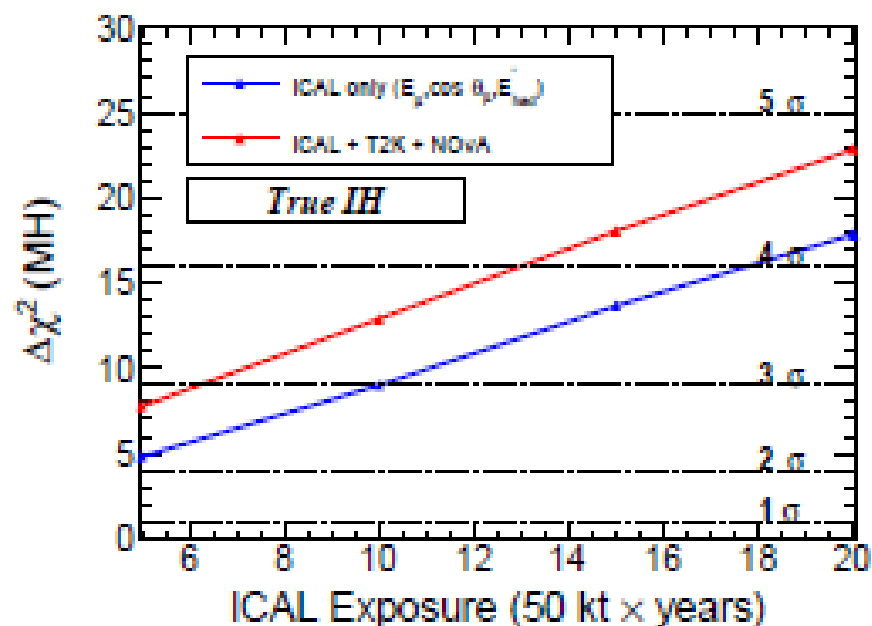
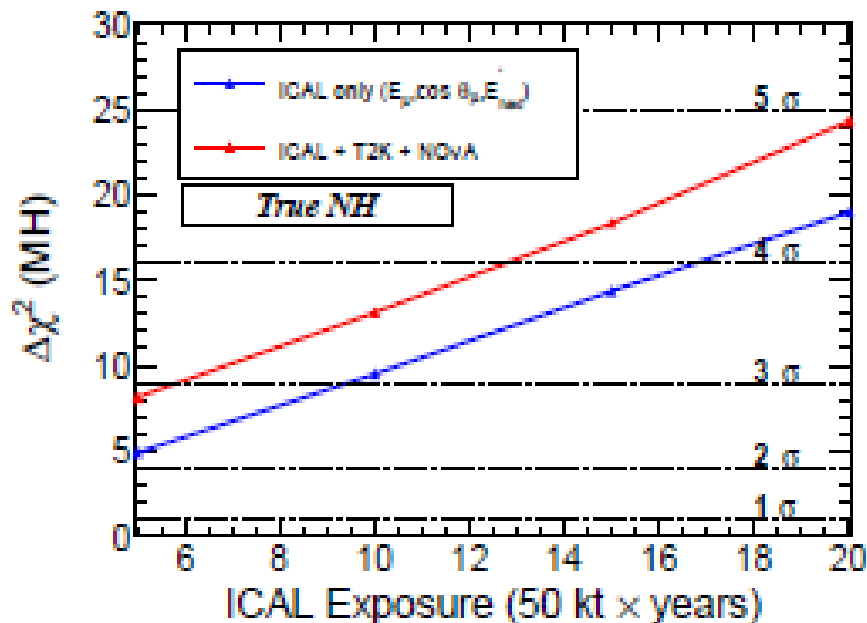


Figure 4: Preliminary results on the hierarchy sensitivity with input normal (left) and inverted (right) hierarchy when ICAL data is combined with the data from T2K (total luminosity of  $8 \times 10^{21}$  pot in neutrino mode) and NO $\nu$ A (3 years running in neutrino mode and 3 years in antineutrino mode) [17].

## Determination of the CP phase

Though ICAL itself is rather insensitive to  $\delta_{\text{CP}}$ , data from ICAL can still improve the determination of  $\delta_{\text{CP}}$  itself, by providing input on mass hierarchy. This is especially crucial in the range  $0 \leq \delta_{\text{CP}} \leq \pi$ , precisely where the ICAL data would also improve the hierarchy discrimination of NO $\nu$ A and other experiments [19].

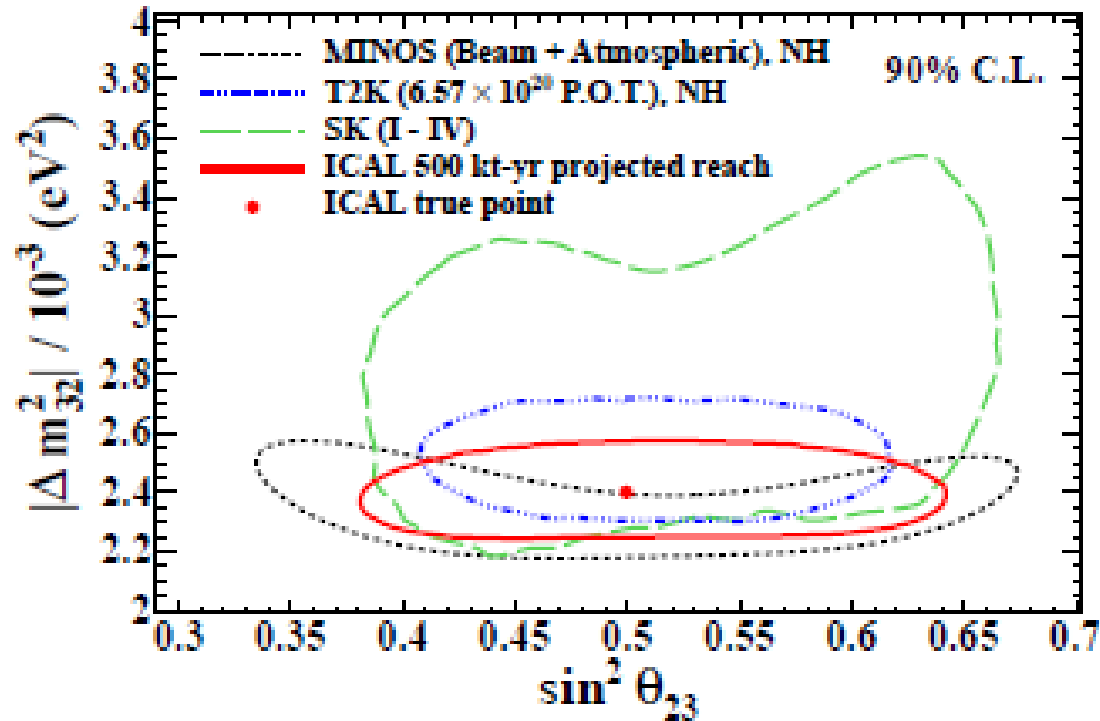


Figure 3: The precision reach of ICAL in the  $\sin^2 \theta_{23} - \Delta m_{32}^2$  plane, in comparison with other current and planned experiments [14]. Information on hadron energy has been included.

With a non-maximal true value of  $\theta_{23}$ , the bounds on  $\sin^2 \theta_{23}$  range would be asymmetric about 0.5. Figure 5.18 shows the sensitivity of ICAL for  $\sin^2 2\theta_{23} = 0.93$  (i.e.  $\sin^2 \theta_{23} = 0.37, 0.63$ ). It may be observed that for  $\theta_{23}$  in the lower octant, the maximal mixing can be ruled out with 99% C.L. with 500 kt-yr of ICAL data. However, if  $\theta_{23}$  is closer to the maximal mixing value, or in the higher octant, then the ICAL sensitivity to exclude maximal mixing would be much smaller.

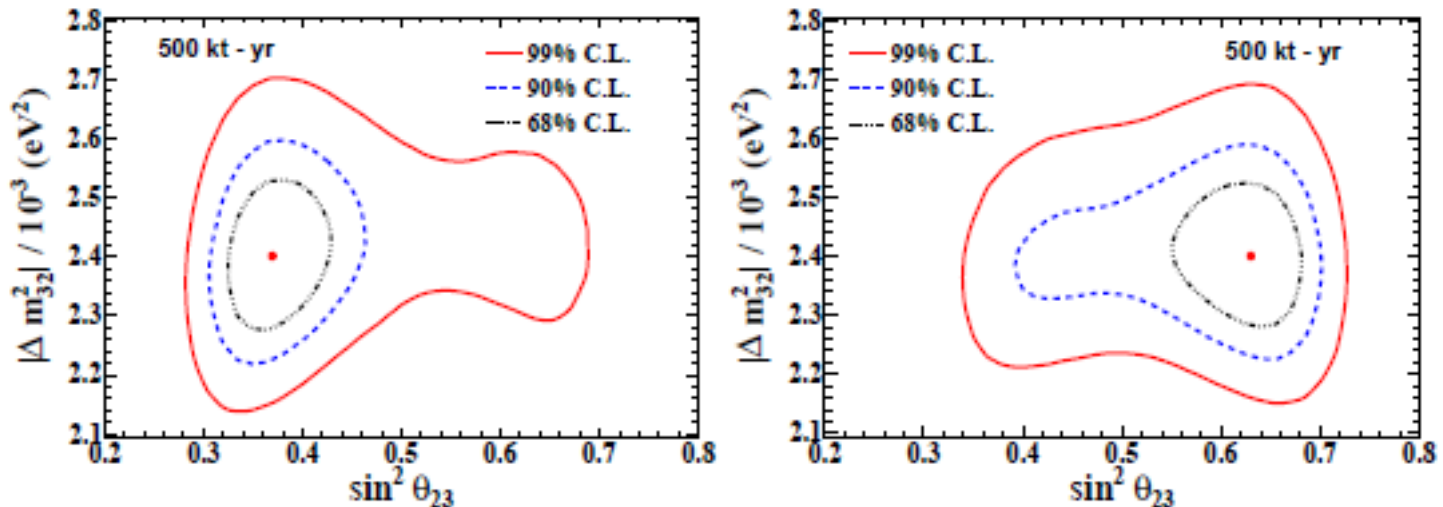


Figure 5.18:  $\Delta\chi^2_{\text{ICAL-PM}}$  confidence level contours in the the  $\sin^2 \theta_{23}$ - $|\Delta m^2_{32}|$  plane, for  $\sin^2 \theta_{23}(\text{true}) = 0.37$  (left panel) and  $\sin^2 \theta_{23}(\text{true}) = 0.63$  (right panel), using the hadron energy information. Here the true hierarchy is NH [14].

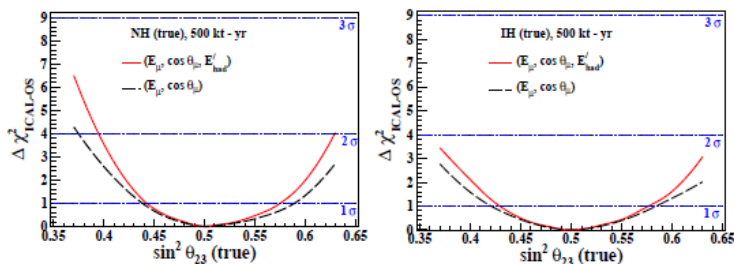


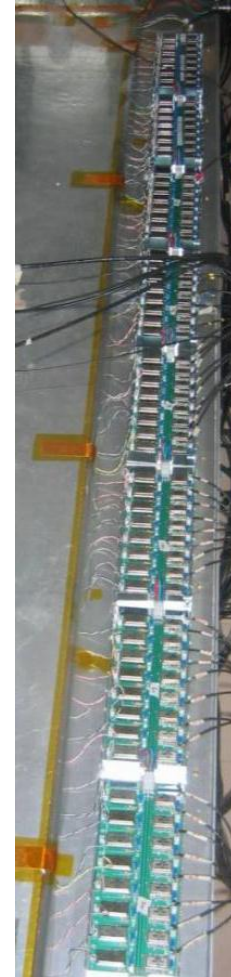
Figure 5.19:  $\Delta\chi^2_{\text{ICAL-OS}}$  for octant discovery potential as a function of true  $\sin^2 \theta_{23}$ . The left panel (right panel) assumes NH (IH) as true hierarchy. The line labelled  $(E_\mu, \cos \theta_\mu)$  denotes results without including hadron information, while the line labelled  $(E_\mu, \cos \theta_\mu, E'_{\text{had}})$  denotes improved results after including hadron energy information. ICAL exposure of 500 kt-yr is considered [14].

# *RPC Development work at TIFR*



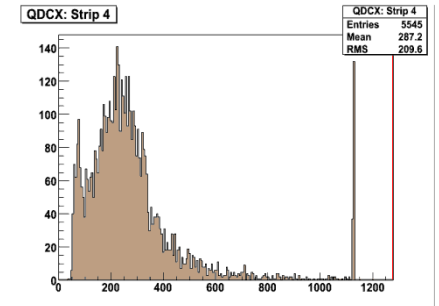
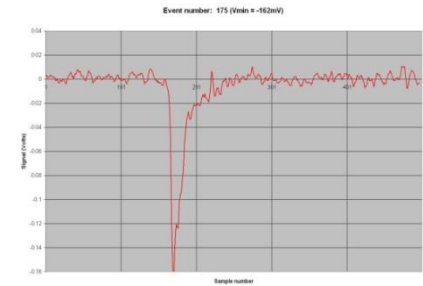
Developing tools for large scale RPC production

# *2m X 2m RPC Test Stand at TIFR*

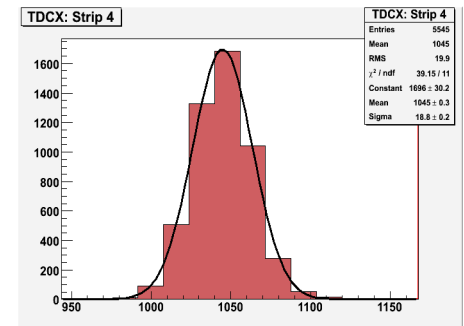


# Testing of RPCs using cosmic ray muons

Gas mixing unit



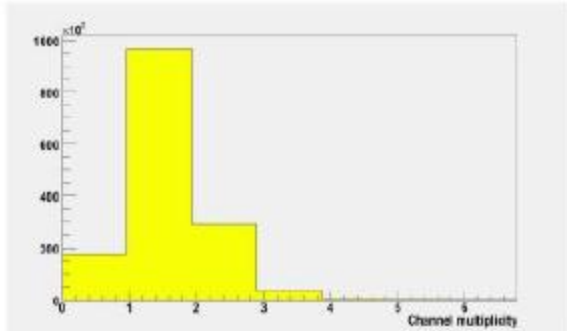
Charge distribution



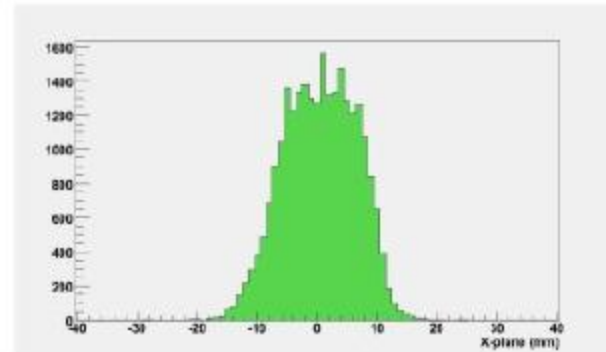
Timing distribution

Streamer mode: HF134a: A : Isobutane:: 62:30:8  
Avalanche mode: A : Isobutane: SF6::95.5:4.5:0.3

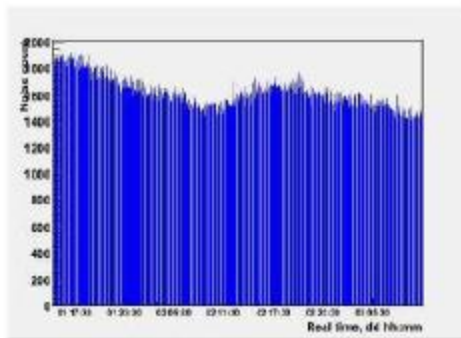
# *RPC performance with cosmic rays*



Strip Multiplicity due to crossing muons



Track residue in mm



Strip noise rate vs time

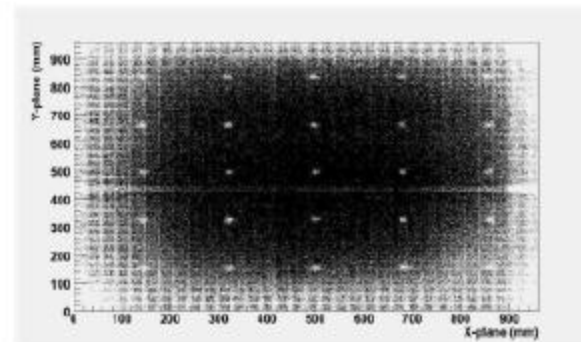
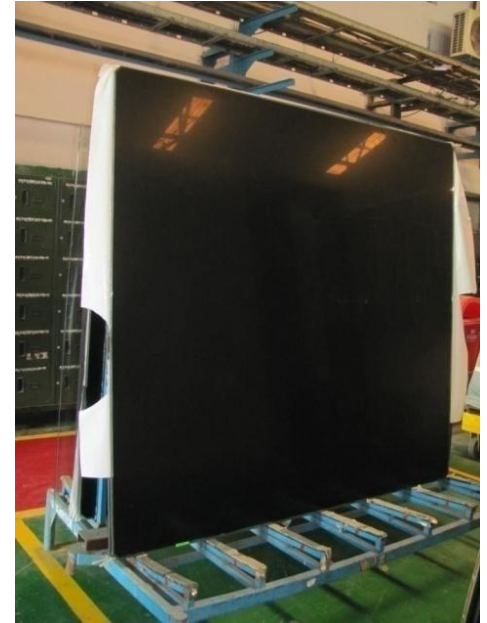
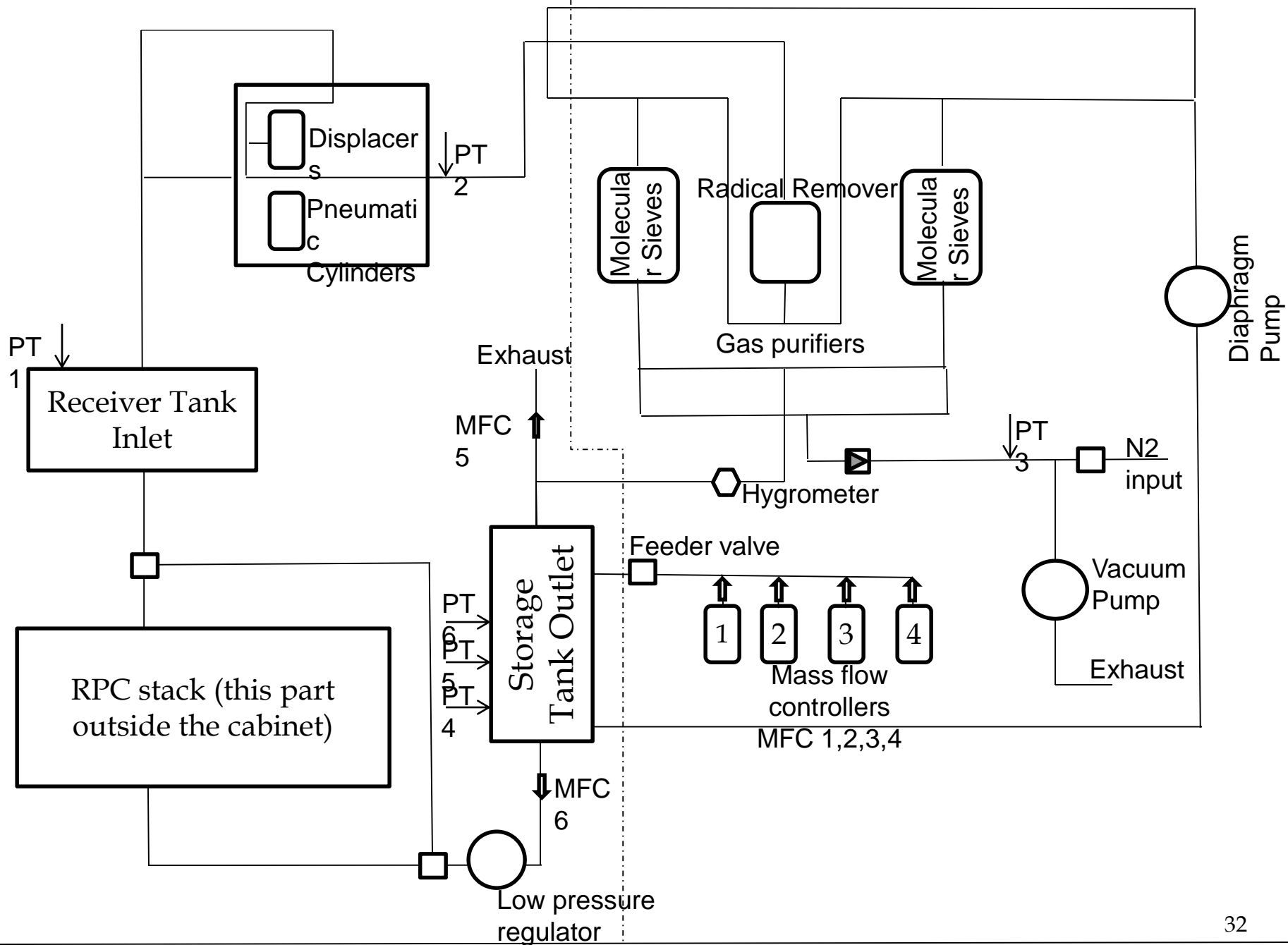


Image of a RPC using muons

# *RPC Fabrication at Asahi Float Glass Ltd. Mumbai*







# *Newly developed gas recirculation system*



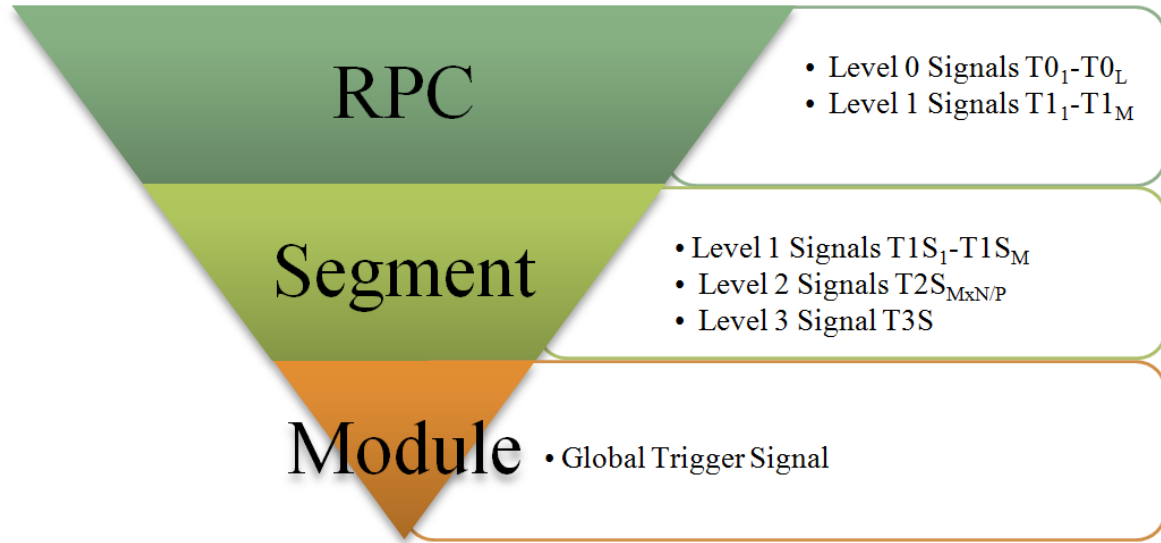
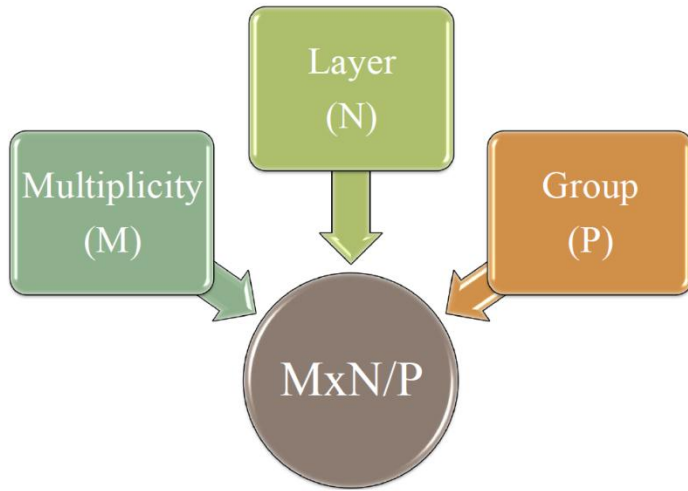
## Gas Recirculation System:

- *Prototype close loop gas recirculation system is under test at TIFR RPC lab.*

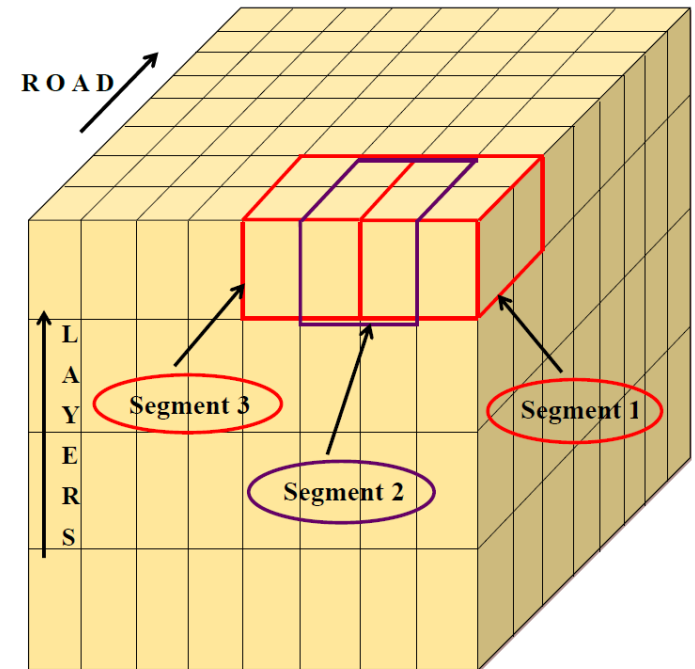
# *Electronics/Daq for RPC system*

- *An in-house front end amp-disc ASIC front end has been designed by the BARC electronics team and fabricated by Euro Practice IC Services. Tested the first version and a revised version has now been fabricated. It is under test.*
- *Overall electronics and DAQ architecture under discussion. Designate RPC as the minimum stand alone unit.*
- *High performance FPGA (featuring a  $\mu$ C softcore), TDC and waveform sampler based RPC-DAQ board is being prototyped.*
- *Data network architecture and hardware being designed to transmit digitized data to the back-end using RPC-DAQ network interface.*
- *Integration of electronics and DAQ hardware with the RPC detector is being finalised. The entire ICAL detector becomes a large Ethernet LAN, suitably segmented, with RPC units as LAN hosts together with the back end DAQ computers*
- *1<sup>st</sup> version of ICAL trigger scheme in place. Many implementation aspects tested, integration issues being addressed.*

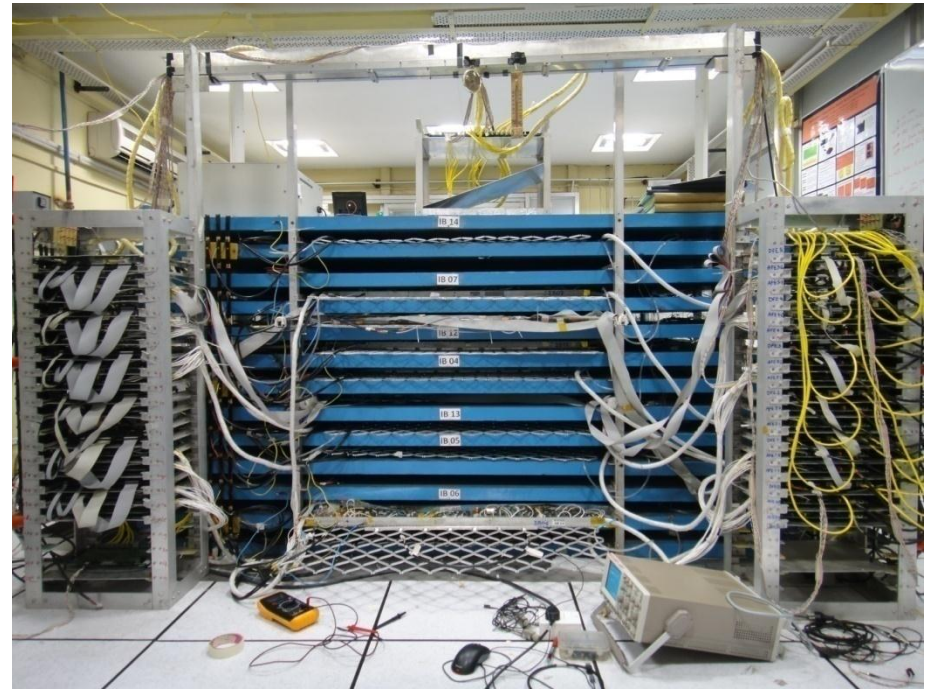
# ICAL Trigger Scheme



- Trigger criteria based on event topology alone.
- Distributed and hierarchical architecture.
- Detector module segmented to generate local trigger.
- Combination of local triggers produces global trigger.
- Global trigger latches event data.

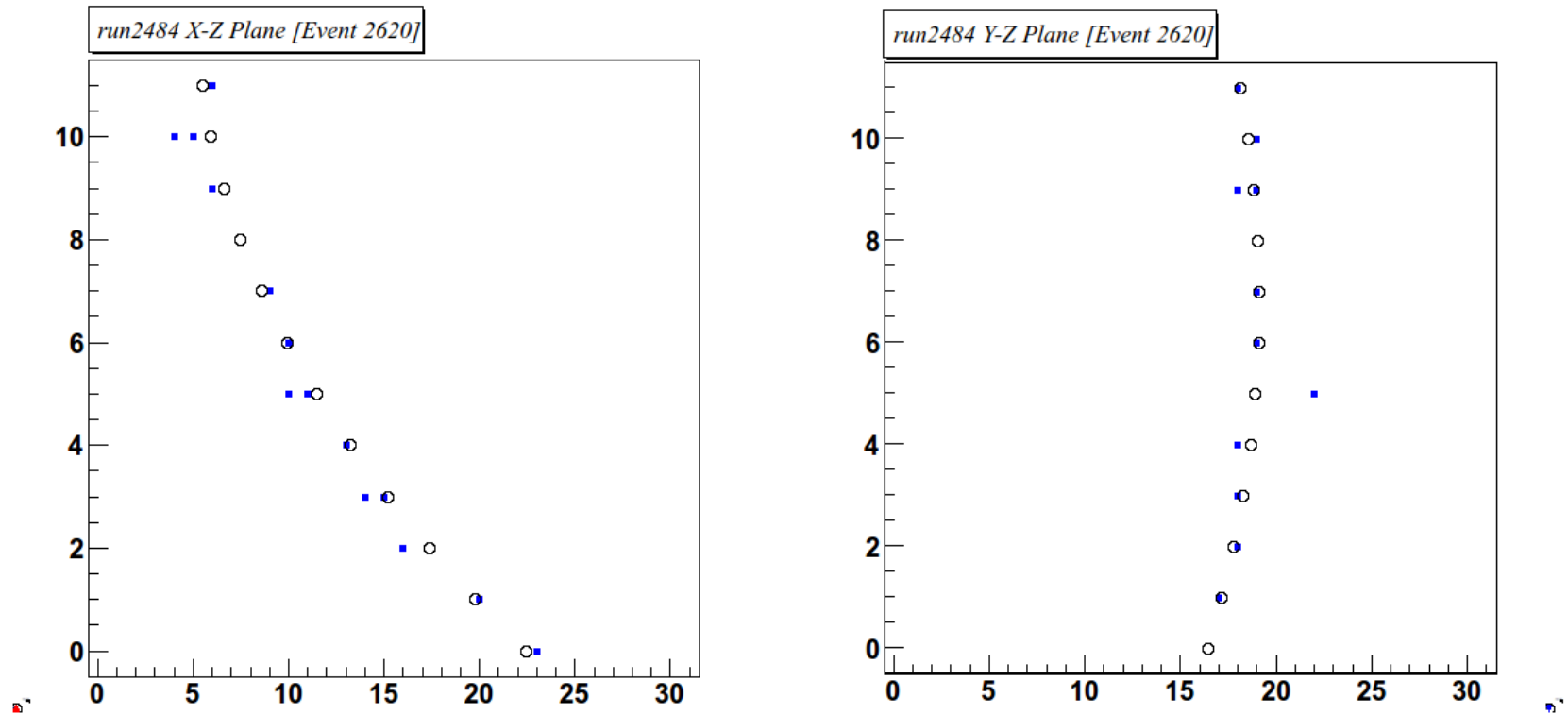


# *INO Prototype magnet at VECC*



- 35 Ton prototype with 12 gaps to house 1mX1m RPCs
- Long term operational experience
- Operate both glass & bakelite RPCs
- Muon track reconstruction with & without magnetic fields.
- Stability & suitability of LV & HV electronics.
- Lab environmental condition studies and student training
- Will be taken to CERN for beam tests and calibration

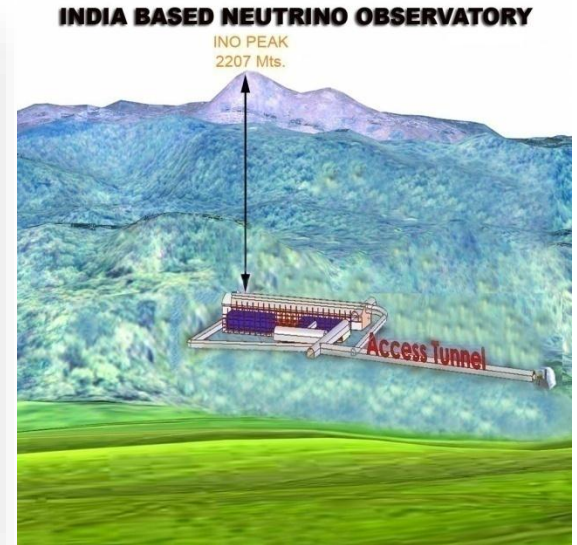
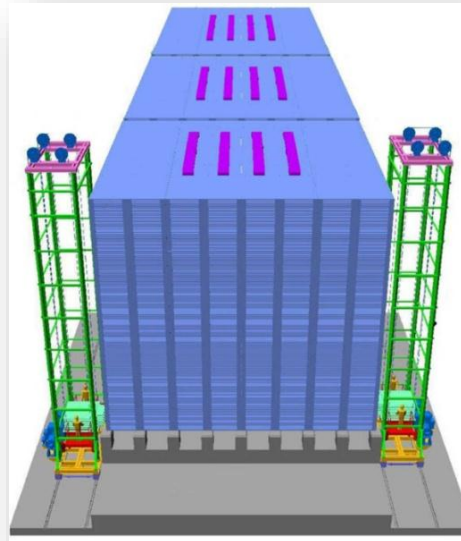
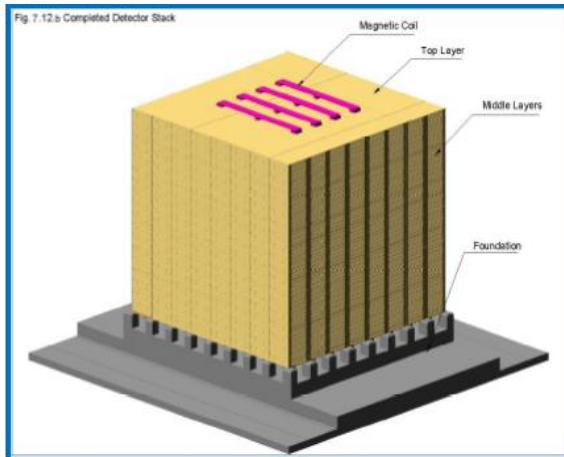
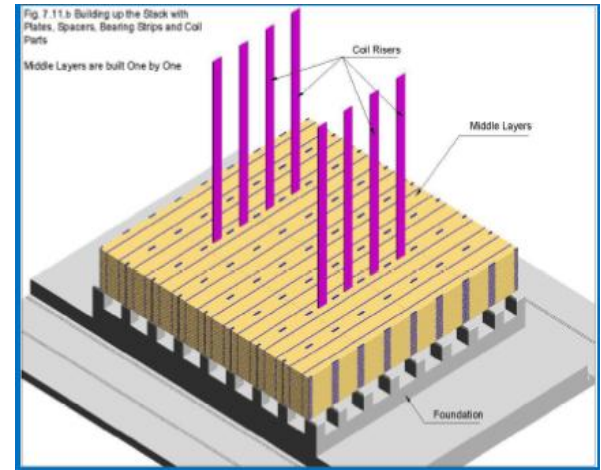
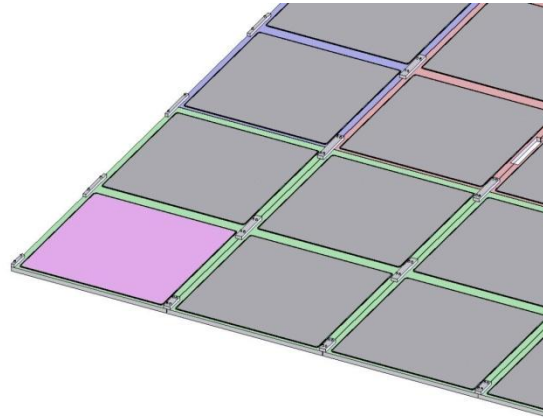
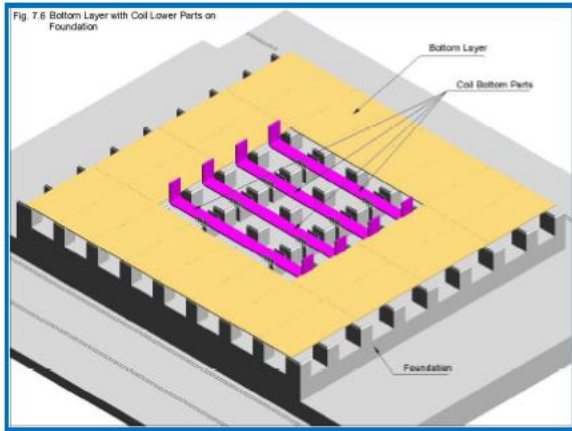
# Cosmic muon tracks seen in the prototype



Hits and fitted points on a bent track for data taken with magnetic field ON condition

Next: 800 ton Engineering proto-type module (8mX8mx20 layers; 400 RPCs).

# Construction of ICAL Magnet



Concepts from DPR prepared by TCE

# Summary - Current Status

Pre-project activities were started with an initial grant of USD ~ 10 Million

Infrastructure development of the Site

Development of IICHEP at Madurai (110 km from INO lab)

Detector R & D is now complete; DPR for Detector is ready

DPR for DAQ system is getting ready

Industrial production of RPCs and associated electronics to start soon

Geo-technical survey for construction of tunnel & cavern to start

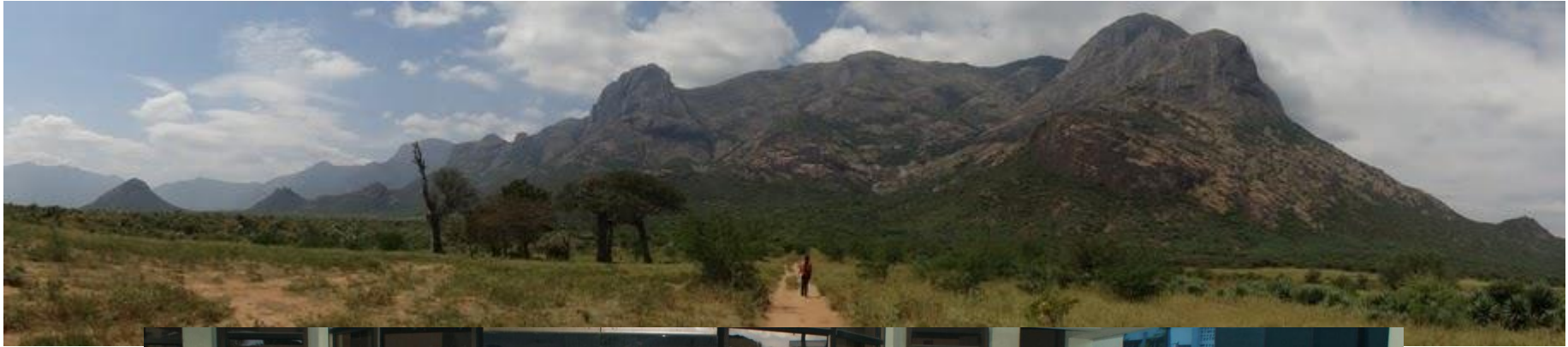
Construction of engineering prototype module at Madurai (in surface lab),  
to test all aspects of construction of detector, efficiency of industrially  
produced RPCs apart from physics studies, logistics of operation.

Full project was approved the funding agencies, (~250M\$) got the clearance  
in the beginning of this year (2015) from the central government for disbursing  
the funds and start construction

We have a strong physics programme , all eager to begin.  
Hopefully, by 2020 detector should be operational at the site.



# *Thank You*



**Collaborators are welcome**  
**<http://www.ino.tifr.res.in/ino>**